# 2. LOWER CRETACEOUS VOLCANIC ROCKS ON CONTINENTAL MARGINS AND THEIR **RELATIONSHIP TO THE KERGUELEN PLATEAU<sup>1</sup>**

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### ABSTRACT

Widespread Lower Cretaceous magmatism occurred along the Indian-Australian/Antarctic margins, and in the juvenile Indian Ocean, during the rifting of eastern Gondwana. The formation of this magmatic province probably began around 120-130 Ma with the eruption of basalts on the Naturaliste Plateau and at Bunbury, western Australia. On the northeast margin of India, activity began around 117 Ma with the Rajmahal continental basalts and associated lamprophyre intrusions. The formation of the Kerguelen Plateau in the Indian Ocean began no later than 114 Ma. Ultramafic lamprophyres (alnoites) were emplaced in the Prince Charles Mountains near the Antarctic continental margin at  $\sim 110$  Ma. These events are considered to be related to a major mantle plume, the remnant of which is situated beneath the region of Kerguelen and Heard islands at the present day.

Geochemical data are presented for each of these volcanic suites and are indicative of complex interactions between asthenosphere-derived magmas and the continental lithosphere. Kerguelen Plateau basalts have Sr and Nd isotopic compositions lying outside the field for Indian Ocean mid-ocean ridge basalts (MORB) but, with the exception of Site 738 at the southern end of the plateau, within the range of more recent hotspot basalts from Kerguelen and Heard Islands. However, a number of the plateau tholeiites are characterized by lower 206Pb/204Pb ratios than are basalts from Kerguelen Island, and many also have anomalously high La/Nb ratios. These features suggest that the source of the Kerguelen Plateau basalts suffered contamination by components derived from the Gondwana continental lithosphere. An extreme expression of this lithospheric signature is shown by a tholeiite from Site 738, suggesting that the southernmost part of the Kerguelen Plateau may be underlain by continental crust.

The Rajmahal tholeiites mostly fall into two distinct geochemical groups. Some Group I tholeiites have Sr and Nd isotopic compositions and incompatible element abundances, similar to Kerguelen Plateau tholeiites from Sites 749 and 750, indicating that the Kerguelen-Heard mantle plume may have directly furnished Rajmahal volcanism. However, their elevated <sup>207</sup>Pb/<sup>204</sup>Pb ratios indicate that these magmas did not totally escape contamination by continental lithosphere. In contrast to the Group I tholeiites, significant contamination is suggested for Group II Rajmahal tholeiites, on the basis of incompatible element abundances and isotopic compositions.

The Naturaliste Plateau and the Bunbury Basalt samples show varying degrees of enrichment in incompatible elements over normal MORB. The Naturaliste Plateau samples (and Bunbury Basalt) have high La/Nb ratios, a feature not inconsistent with the notion that the plateau may consist of stretched continental lithosphere, near the ocean-continent divide.

## INTRODUCTION

Kerguelen Plateau and Broken Ridge, which were originally contiguous, have an estimated combined crustal volume of approximately 57  $\times$  10<sup>6</sup> km<sup>3</sup>, comparable to the Ontong Java Plateau in the Pacific (Schubert and Sandwell, 1989). The formation of this major volcanic edifice in the juvenile Lower Cretaceous Indian Ocean was accompanied by magmatism

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along the continental margins of eastern Gondwana, proximal to the plateau. Specifically, continental basalts were erupted in northeastern India in the Rajmahal region and in western Australia at Bunbury and, offshore, on the Naturaliste Plateau. Lamprophyre intrusions occur in India and Antarctica. These occurrences are summarized in Figure 1.

One of the main objectives of Ocean Drilling Program (ODP) Legs 119 and 120 was to recover basement material that would help test the various hypotheses that have been proposed for the origin of the Kerguelen Plateau. However, any model for the origin of the plateau should also account for the associated continental magmatism. Previously proposed models for the origin of the Kerguelen Plateau include:

1. Continental fragment: Dietz and Holden (1970) suggested that the plateau contains rifted continental crust, a remnant of the Gondwana supercontinent. Schlich et al. (1971) considered the sedimentary section of the northern plateau to have continental affinities whereas Ramsay et al. (1986) suggested that the Southern Kerguelen Plateau, based on the granitic material and metasediments in free-fall grab samples, contains continental crust.

2. Uplifted oceanic crust: Houtz et al. (1977) suggested that the Kerguelen Plateau was, in part, a remnant of Cretaceous ocean basin crust that formed to the west of Australia during the separation of India from Australia-Antarctica.

<sup>&</sup>lt;sup>1</sup> Wise, S. W., Jr., Schlich, R., et al., 1992. Proc. ODP, Sci. Results, 120: College Station, TX (Ocean Drilling Program).



Figure 1. Reconstruction of eastern Gondwana at ca. 120 Ma showing the location and nature of continental Lower Cretaceous volcanism immediately before, or contemporaneous with, the formation of the Kerguelen Plateau at ca. 110-115 Ma (modified from Kent, 1991). Also shown are the position of rift zones that acted as major drainage pathways for up to 150 m.y. before the beginning of igneous activity (Kent, 1991). Inset shows (A) the present-day location of the Kerguelen-Heard plume (KHP) between Kerguelen and Heard islands and (B) its position at 110 Ma, at the time of formation of the Kerguelen Plateau. RT = Rajmahal Traps; BB = Bunbury Basalt; and PCM = Prince Charles Mountains. It has been suggested (e.g., Davies et al., 1989; Storey et al., 1989) that this entire magmatic province may be attributable to the KHP, which has more recently produced Dupal-type oceanic basalts (Hart, 1984) on Kerguelen and Heard islands in the southern Indian Ocean (Storey et al., 1988; Barling and Goldstein, 1990); Gautier et al., 1990).

3. Hotspot related: A hotspot or mantle plume origin for the Kerguelen Plateau has been suggested by many authors based on plate reconstructions and geochemistry (e.g., Luyendyck and Rennick, 1977; Duncan, 1978; Peirce, 1978; Morgan, 1981; Storey et al., 1989; Weis et al., 1989; Davies et al., 1989).

This paper represents a preliminary report of work in progress in which we review the setting of the Kerguelen Plateau and associated continental volcanism, present new major and trace element data for each of the volcanic occurrences, and examine the data in terms of the above models.

### **GEOLOGICAL SETTING**

### Kerguelen Plateau

The Kerguelen Plateau, located on the Antarctic Plate between  $46^{\circ}$  and  $64^{\circ}$ S, is one of the most important physiographic features on the Indian Ocean floor (Fig. 2). It runs north-northwest for approximately 2300 km, is about 500 km wide, and rises over 3 km above the surrounding ocean floor. It is separated from the Antarctic continent by the 100km-wide, 3500-m-deep Princess Elizabeth trough. Seafloor spreading at the Southeast Indian Ridge (SEIR) separated the Kerguelen Plateau and Broken Ridge approximately 45-42 Ma ago (e.g., Houtz et al., 1977; Munschy and Schlich, 1987).

The Kerguelen Plateau has been divided into northern and southern sectors by Schlich (1975) and Houtz et al. (1977). The northern part generally lies in water depths of less than 1000 m and contains the Eocene to Quaternary volcanic islands of Kerguelen, McDonald, and Heard. The boundary between the northern and southern sectors of the plateau lies immediately south of Heard Island. The transition zone exhibits a complex bathymetry with a large west-trending spur, the Elan Bank, extending from the main plateau over a distance of some 600 km. The southern portion of the plateau is generally deeper, with water depths between 1000 and 3000 m, and has a more subdued topography. It consists of a broad anticlinal arch affected by multiple stages of normal faulting that result in horst and graben development (e.g., Coffin et al., 1986). The axis of this arch is marked by the 77°E Graben.

Seismic refraction studies on and in the vicinity of the Kerguelen Archipelago by Recq et al. (1983), Recq and Charvis (1986), and Recq et al. (1990) suggest a crustal thickness of between 15 and 23 km for this part of the plateau.



Figure 2. Bathymetric map (in meters) of the Kerguelen Plateau showing the location of ODP drill sites (circles; filled circles denote basement sites) and MD48 dredge stations (stars).

These authors note that the seismic velocity vs. depth profile is typical of oceanic islands. Houtz et al. (1977) inferred a crustal thickness of 20–23 km for the Southern Kerguelen Plateau based on the gravity data.

According to Schlich (1982), the oldest magnetic anomaly bordering the plateau is 34 (equivalent to 84 Ma, following the time scale of Kent and Gradstein, 1985), which occurs in the Crozet Basin. However, a recent revision of the magnetic profiles in the Enderby Basin on the southwestern margin of the plateau (Li, 1988) has indicated the presence of magnetic anomalies M1–M7 (122–128 Ma).

The first drill samples from the Kerguelen Plateau basement were obtained on the central and southern parts of the plateau during ODP Legs 119 and 120 (Fig. 2). In addition, dredge samples (including basalt) from the 77°E Graben and the northern part of the plateau were obtained by the Marion Dufresne in 1986 (Cruise MD48; Leclaire et al., 1987). Detailed descriptions of the Cruise MD48 dredge samples and the Leg 119 and 120 cores are given by Bassias et al. (1987), Barron, Larsen, et al. (1989), and Schlich, Wise, et al. (1989), respectively. Some of the Kerguelen Plateau basement basalts appear to have been erupted in a subaerial environment, as evidenced by the presence of oxidized flow tops at Site 747 and overlying nonmarine sediments, of probable Albian age, at Site 750 (Schlich, Wise, et al., 1989). The age of the overlying sediment at the remaining sites is pre-Turonian (limestone) at Site 738, lower Santonian (nannofossil chalk) at Site 747, and early Eocene (nannofossil ooze) at Site 749. In addition, at Site 748, some 150-200 m above the seismic basement, a highly altered alkali-basalt flow was drilled, overlain by an undated conglomerate and then by upper Albian-Turonian sandstones (Schlich, Wise, et al., 1989).

Radiometric dating of the Kerguelen Plateau basement has so far been restricted to a few samples from the centralsouthern parts of the plateau (Leclaire et al., 1987; Whitechurch et al., this volume). Whole-rock dating on the leastaltered Leg 120 basalts give  ${}^{40}\text{Ar}/{}^{39}\text{Ar}$  "plateau ages" of 109.2  $\pm$  0.7 Ma for Site 749 located on the Banzare Bank and 118.2  $\pm$  5 Ma for Site 750 in the Raggatt Basin (Whitechurch et al., this volume). The Leg 120 results show a close correspondence to a plagioclase K-Ar age of 114  $\pm$  1 Ma from a dredged tholeiite (dredge station 5) on the 77°E Graben (Leclaire et al., 1987). However, all of the three dated samples are within 400 km of each other (Fig. 2). These dates, when coupled with the age of the oldest magnetic anomalies in the Crozet Basin, suggest that the plateau may span at least 30 m.y. Further sampling and dating studies are required, particularly on the northern and southern extremities of the plateau.

## Naturaliste Plateau and Bunbury Basalt

The Naturaliste Plateau (Fig. 3) is a submarine topographic high rising about 2500 m above the surrounding ocean floor at the southwestern corner of continental Australia, and separated from it by a deep channel, the Naturaliste Trough (Jongsma and Petkovic, 1977). The plateau runs approximately 400 km east-west and 250 km north-south. It dips slightly to the north with steep flanks to the south and west. On the northern flank a regular continental rise apron has developed east of 111°E, continuous with that of the margin of Western Australia, but the western part of the northern flank is steep. To the northwest, the Naturaliste Plateau abuts oceanic crust of Lower Cretaceous age M4 (126.5 Ma) whereas M11 (132.5 Ma) represents the onset of seafloor spreading in the Perth Basin.

The Deep Sea Drilling Program drilled on the Naturaliste Plateau at Sites 258 (Leg 26) and 264 (Leg 28). Site 258 consisted of some 525 m of marine Cenozoic and Cretaceous sediments. The lowest part of the section consisted of glauconitic sandstone and silty clay, overlain by 251 m of mid-Albian to Cenomanian ferruginous clays. Most of the detrital constituents of this lower part were probably derived from the erosion of basaltic volcanic rocks, the source probably being an elevated, subaerially exposed part of the plateau (Davies et al., 1974). At Site 264, 171 m of marine carbonate sediments were drilled, ranging in age from Holocene to Santonian. The drill bottomed after a further 35 m of altered, nonfossiliferous volcaniclastic conglomerate, estimated by Hayes, Frakes, et al. (1975) to be Cenomanian or older in age. Coleman et al. (1982) consider that the top of this conglomerate corresponds to a prominent reflector that Jongsma and Petkovic (1977) described and which they were able to trace over much of the subsurface of the plateau. Dredge samples recovered from the steep northwestern flanks of the plateau by the Eltanin in 1972 were reported by Heezen and Tharp (1973) as containing rocks of continental origin, although reexamination of the dredge haul by Coleman et al. (1982) revealed that the rock clasts were actually altered tholeiitic basalts.

The Bunbury Basalt represents a possible onland correlative to the Naturaliste Plateau. It consists of a few thick tholeiitic flows that crop out over 140 km of the southern portion of the Perth basin (Fig. 3). K-Ar ages for the Bunbury Basalt range from 88 to 105 Ma (McDougall and Wellman, 1976). However, a dolerite from the Perth Basin has been dated at  $136 \pm 3$  Ma (Playford et al., 1976), and a similar age has been inferred for the Bunbury basalt (e.g., Davies et al., 1989). More precise dating of both the Naturaliste Plateau and the Bunbury Basalt are required.

### **Rajmahal Volcanic Formation**

The Rajmahal Volcanic Formation is exposed on the continental margin of northeast India, cropping out over an



Figure 3. Bathymetric map (in meters) of the southwest Australian continental margin (after Coleman et al., 1982) showing the location of the Naturaliste Plateau, major fracture zones, seafloor magnetic anomalies, and the location of *Elianin* dredge and Bunbury Basalt samples. Also shown are DSDP Sites 257, 258, and 264. Y = Yilgarn Block (Archean), L = Leeuwin Block (Proterozoic), and BT = Bunbury trough. Inset shows location of other major plateaus and basins on the west Australian continental margin: NP = Naturaliste Plateau, PB = Perth Basin, WP = Wallaby Plateau, Z = Zenith Seamounts, CB = Cuvier Basin, EP = Exmouth Plateau, AB = Argo Basin, and SP = Scott Plateau.

area of around 4,300 km<sup>2</sup> in the Rajmahal Hills, Bihar (Fig. 4). The formation consists dominantly of quartz-normative tholeiites, with minor olivine-basalts and basaltic andesites, totaling some 200–300 m in thickness. This wedge of subaerial volcanics overlies a thin sequence of Early Permian to Early Cretaceous sediments (the Gondwana Supergroup), in turn resting on a high-grade Precambrian basement complex (the Chotanagpur mobile belt). The volcanics are interbedded with thin sedimentary or volcaniclastic horizons, containing locally abundant Lower Cretaceous plant fragments (Sengupta, 1988). Dating of the basalts by  $4^{0}$ Ar/<sup>39</sup>Ar (Baksi, 1989) also supports a Early Cretaceous age of 117 ± 1 Ma. These rocks are believed to be cogenetic with a suite of tholeiites and alkali basalts exposed along the Dauki lineament near Sylhet, Meghalaya. In this area, a  $60 \times 40$  km strip of lavas, some 500–600 m thick (?), has been documented by Talukdar (1967) and Talukdar and Murthy (1970). The volcanic succession comprises alternating tholeiitic and felsic pyroclastic/rhyolite units, overlain by nepheline tephrites, dipping 19°–25° toward the southeast. The sequence is cut by a basic dyke swarm, the extent and nature of which is uncertain at present. Early Cretaceous alkaline complexes also occur in this region (Chattopadhyay and Hashimi, 1984).



Figure 4. Simplified geological map of the Rajmahal region showing the locations of the basalts analyzed.

37

During the 1950s, the subsurface geology of the Bengal Basin was the focus of an integrated hydrocarbon prospect by the Standard Vacuum Oil Company (Mobil/Esso). Seismic reflection profiles and borehole data revealed the presence of a seaward-dipping reflector sequence on the continental margin, comprising a thin wedge of mafic flows (Biswas, 1963; Sengupta, 1966). These flows form the tops of tilted fault blocks, similar in structural style to those on the opposing continental margin of southwest Australia (Marshall and Lee, 1989), and suggestive of episodic extension and, by inference, transtension. Kent (1991) relates these features to episodic doming and extension of eastern Gondwana lithosphere from Early Permian to Early Cretaceous times.

### **Damodar Valley and Darjeeling Lamprophyres**

Recent work by Sarkar et al. (1980) and Paul and Potts (1981) provided the first geochemical and geochronological data on Indian Cretaceous lamprophyres. These rocks intrude sediments of the Gondwana Supergroup in the Damodar and Darjeeling coal fields and are also found along lineament zones in the Proterozoic metamorphic complexes of the Shillong Plateau (Fig. 1). They consist of both alkaline and ultramafic varieties, emplaced in the form of horizontal cylinders or, more rarely, as thin dykes and sills. K-Ar ages on biotite phenocrysts and whole-rock samples range from 121 to 105 Ma (Sarkar et al., 1980). It is presently uncertain as to whether these intrusions postdate or are contemporaneous with the eruption of the Rajmahal tholeiites.

### **Prince Charles Mountains Lamprophyres**

At the southwest corner of Radok Lake ( $70^{\circ}54'S$ ,  $67^{\circ}56'E$ ) in the Prince Charles Mountains of Antarctica (Fig. 1), two sills occur in conglomerate near the base of an upper Permian sedimentary sequence (Kemp, 1969). They are examples of the fairly rare ultramafic lamprophyre alnoite and have been described by Walker and Mond (1971). The lower sill is about 4 m thick at its western extremity, thickening eastward to 20 m. The upper sill, which is stratigraphically about 300 m above the lower, is about 5 m thick. The sills are exposed in vertical cliffs and can be traced horizontally over a distance of 1 km. Thin dykes also occur in the sediments 6.5–11 km northeast of the sills. K-Ar dating on phlogopite separates from the sills gave ages between 108 and 110 Ma (Walker and Mond, 1971).

# MAJOR AND TRACE ELEMENT GEOCHEMISTRY

#### Methods

Samples were cleaned before being crushed in an agate shatterbox. Trace element analysis was conducted on 46-mmdiameter powder briquettes, made from 15 g of rock powder by adding 20 drops of a 7% solution of polyvinyl alcohol and subjecting the mixture to a pressure of 15 tons in a steel die. For major element analysis, the rock powders were dried overnight at 120°C before igniting them at 800°C. This was followed by accurately weighing 1 g of ignited rock powder with 5 g of Johnson Matthey spectroflux 100B into a platinum crucible and fusing the mixture at 1200°C. The melt was pressed into a glass disk on an aluminum mold.

X-ray fluorescence (XRF) analysis was conducted at the University of Leicester using a Philips PW1400 automatic spectrometer. Major elements and Nb, Zr, Y, Sr, Rb, Th, and Ni were determined using a Rh X-ray tube. The trace elements V, Cr, Ba, La, Ce, and Nd were determined using a W X-ray tube. Details of operating conditions and data quality are given by Marsh et al. (1980) and Weaver et al. (1983).

Instrumental neutron activation analysis (INAA) of the rare-earth elements (REE) and Th, Ta, Hf, and Sc was conducted on selected samples at the University of Durham using a method similar to that detailed by Leat et al. (1990), counting samples for 24 hr.

The XRF analyses of standard W1 and the INAA analyses of standards BR and BOB-1, performed in conjunction with this work, are given in Saunders et al. (in press).

### **Kerguelen Plateau**

Legs 119 and 120 recovered Kerguelen Plateau basement at Sites 738, 747, 749, and 750 (Fig. 2), which consisted of moderately fresh to highly altered basalt. The secondary mineral assemblage consisted of smectites, zeolites, and calcite. The basalts range from aphyric to highly porphyritic, containing phenocrysts of plagioclase, clinopyroxene and olivine. The lava flow drilled at Site 748 is a highly altered analcite-bearing alkali basalt containing plagioclase and clinopyroxene phenocrysts. The MD48 basement dredge samples recovered from the 77°E Graben and the northwest flanks of the Kerguelen Plateau consisted of moderately to highly altered basalt (Bassias et al., 1987; Leclaire et al., 1987; Davies et al., 1989).

New major and trace element analyses for Leg 120 and MD48 Kerguelen Plateau basalts are given in Table 1. Most are olivine normative tholeiites, with MgO, Ni, and Cr contents for the Leg 120 samples ranging from 5.8% to 10.2%, 29 to 156 ppm, and 83 to 401 ppm, respectively. The MD48 dredge samples contain more evolved compositions, with MgO varying between 3.9% to 5.7%. The tholeiites have been variably affected by hydrothermal alteration, as indicated by their large range in volatile contents (e.g., LOI = 0.1% to 7.3%) and the generally nonsystematic behavior of the more mobile elements (e.g., Rb, K, and Na) with magmatic fractionation indexes such as MgO or Zr (Schlich, Wise, et al., 1989).

In terms of immobile incompatible element abundances and ratios, Leg 120 basalts exhibit significant inter- and intrasite variations. For example, Site 747 basalts have lower Ti/Zr ratios (74 to 83) than those of Site 749 (94 to 139) and Site 750 (150 to 185), suggesting that the first are derived from a more enriched source (Fig. 5). The Site 749 tholeiites appear to consist of two distinct magma types: most of the data have Ti/Zr ratios clustering around 100, but a second, more depleted, group has Ti/Zr ratios ranging from 122 to 139. The MD48 tholeiites range in Ti/Zr ratios from 74 to 97.

Figure 6 consists of MORB-normalized multi-element and chondrite-normalized REE plots and illustrates the large range in incompatible element ratios exhibited by Kerguelen Plateau tholeiites. Moderate to strong enrichment of Th and the light REE (LREE) is shown by samples from Site 747 ( $La_N/Yb_N =$ 3.2 to 4.7), Site 738 (La<sub>N</sub>/Yb<sub>N</sub> = 4.2), and the MD48 basalts dredged from the 77°E Graben and the northern part of the Kerguelen Plateau (La<sub>N</sub>/Yb<sub>N</sub> = 2.2 to 6.1). In contrast, basalts from Site 749 are slightly enriched in the LREE ( $La_N/Yb_N =$ 1.53 to 1.80), whereas basalts from Site 750 in the Raggatt Basin have the lowest La<sub>N</sub>/Yb<sub>N</sub> ratios (0.88 to 1.12), being the most similar to normal MORB. Plagioclase accumulation is the most likely explanation for the Sr spike in the strongly phyric (30% plagioclase phenocrysts) Sample 120-749C-16H-7, 69-76 cm (Fig. 6). A notable feature of the Kerguelen Plateau tholeiites from Sites 738, 747, 750, and MD48 dredge stations 2, 5, and 6 is their fairly low Nb and Ta concentrations relative to adjacent, similarly incompatible elements (e.g., Th and La) in the multi-element plots (Fig. 6). High La/Nb ratios (>1) are more normally associated with subduction-related magmatism (e.g., Saunders et al., 1980; Gill, 1981) and some continental flood basalts (e.g., Mantovani et al., 1985) and lamprophyres (e.g., Paul and Potts, 1981).

### Naturaliste Plateau and Bunbury Basalt

Analyses of Naturaliste Plateau basalts are given by Coleman et al. (1982), who report major and trace element data on seven dredge samples recovered by the *Eltanin*. They consist of altered plagioclase- and clinopyroxene-phyric basalts. Analyses of the Bunbury Basalt are given by Burgess (1978), averages of which are given by Coleman et al. (1982). The Bunbury Basalt consists of moderately fresh plagioclase- and pyroxene-phyric flows.

New major and trace element data for five *Eltanin* dredge basalts from the Naturaliste Plateau and seven samples of the Bunbury Basalt are given in Table 2. The location of the samples are shown in Figure 3. *Eltanin* Samples EL5512-7 and EL5512-8 were also analyzed by Coleman et al. (1982), and their data are shown in Table 2. The two data sets show good agreement, with the exception of Rb.

In contrast to the Kerguelen Plateau basement, the Naturaliste Plateau basalts are all strongly quartz-normative tholeiites (normative quartz = 12.9 to 15.9). Their large normative quartz content is probably, in part, a consequence of alteration, as is also the presence of normative corundum (Table 2). The Bunbury Basalt samples are weakly quartz-normative tholeiites. In terms of their compatible elements, the Naturaliste Plateau samples are moderately evolved with MgO, Ni, and Cr ranging from 4% to 5.1%, from 111 to 139 ppm, and from 188 to 555 ppm, respectively. The Bunbury Basalt samples have slightly higher MgO contents (5.5% to 6.5%) but lower Ni (37 to 58 ppm) and Cr (39 to 146 ppm) values (Coleman et al., 1982, this volume).

The Naturaliste Plateau basalts show stronger LREE enrichment ( $\text{La}_N/\text{Yb}_N = 3.1$  to 4.7) than do the Bunbury Basalt samples ( $\text{La}_N/\text{Yb}_N = 1.54$  to 2.25) (Fig. 7). Both are enriched in highly incompatible elements (e.g., Th, LREE) relative to normal MORB. A feature of the multi-element plots for both the Naturaliste Plateau and Bunbury Basalt samples are the troughs at Nb and Ta, with La/Nb ratios ranging from 1.4 to 3.3.

## **Rajmahal Basalts and Indian Lamprophyres**

Previous studies on the volcanic rocks of the Rajmahal Traps and Sylhet Traps (e.g., Mahoney et al., 1983; Baksi et al., 1987) have shown that they consist predominantly of quartz-normative tholeiites and basaltic andesites with a phenocryst assemblage of plagioclase, augite, Fe-Ti oxides, and, more rarely, olivine. Analysis of a few samples from the Bengal Basin has also revealed the presence of alkali basalts and olivine-normative tholeiites (Baksi et al., 1987).

Table 3 presents 25 new analyses of the Rajmahal basalts; their locations are shown in Figure 4. They are all quartz-normative tholeiites and exhibit varying contents of MgO, Ni, and Cr, ranging from 5.3% to 8%, from 18 to 147 ppm, and from 105 to 673 ppm, respectively.

With the exception of Sample RB 88/34, the Rajmahal tholeiites define two distinct groups in terms of their incompatible element abundances and ratios (Groups I and II; Figs. 5C, 8, and 9). Group I tholeiites are characterized by higher Ti/Zr (94 to 118), and lower Zr/Y (3.1 to 3.9),  $K_2O$  (0.15% to 0.55%), and Rb (0.3 to 7.2 ppm) than the Group II tholeiites (Ti/Zr = 61 to 73; Zr/Y = 4.4 to 4.8;  $K_2O = 0.56\%$  to 1.2%, Rb = 17 to 33 ppm). The K/Rb ratios of Group I tholeiites (K/Rb = 242 to 1038) overlap with Group II tholeiites (K/Rb = 248 to 422), but are more variable as they range to higher values. The general enrichment in highly incompatible elements (e.g., Th, LREE) of Group II tholeiites over Group I tholeiites is illustrated in Figure 9. It is also interesting that overall Group I tholeiites (La/Nb = 1.5 to 2.4). The La/Nb ratios of

some Group I tholeiites (e.g., La/Nb  $\leq$  1) are typical of oceanic-island basalts.

The geochemistry of the Lower Cretaceous Indian lamprophyres that intrude into the Gondwana sediments has been described by Paul and Potts (1981) and Middlemost et al. (1988). The Indian lamprophyres are characterized by variable SiO<sub>2</sub> (39.0% to 54.6%), moderate to very high K<sub>2</sub>O (2.2% to 5.9%), MgO (4.7% to 16.8%), Ni (30 to 930 ppm), and Cr (31 to 977 ppm) and low to high Al<sub>2</sub>O<sub>3</sub> contents (4.3% to 15.1%). The lamprophyres show variable enrichment in LREE (La<sub>N</sub>/Yb<sub>N</sub> = 29 to 110), while exhibiting relative depletion in Ta (and Nb) (Fig. 9).

#### **Prince Charles Mountains Lamprophyres**

The petrography of the Radok Lake lamprophyres in the Prince Charles Mountains (Fig. 1) has been described by Walker and Mond (1971). The samples show varying degrees of deuteric alteration, the least altered rocks consisting mostly of phlogopite, melilite, olivine, and nepheline with minor amounts of opaque iron minerals, perovskite, and apatite. Secondary minerals make up 30% to 70% of the rock and include serpentine, chlorite, talc, prehnite, calcite, dolomite, and possibly antigorite (Walker and Mond, 1971; N.C.N. Stephenson pers. comm., 1989). Based on their mineralogy and major element chemistry, they classify as alnoites within the ultramafic lamprophyre group (Walker and Mond, 1971).

Two whole-rock samples collected from the lower sill at Radok Lake have been analyzed for major and trace elements (Table 4). They have low  $SiO_2$  (37.2%), high MgO (13.9% and 16.1%) and Ni (239 and 281 ppm). The alnoites are enriched in LREE, Ta, Nb, Th, K, and Rb. Unlike the ultra-potassic Indian lamprophyres, they have low La/Nb ratios (0.58 and 0.75) and other trace element similarities to ocean island basalts (Fig. 10).

#### **ISOTOPE GEOCHEMISTRY**

#### Kerguelen Plateau

Preliminary Sr, Nd, and Pb isotopic data for Kerguelen Plateau and Rajmahal basalts are shown in Figures 11-13; the complete data set will be given elsewhere. Samples from the Kerguelen Plateau show a remarkably large variation in isotopic composition; for example,  $\epsilon_{Nd}$  ranges from 5.2 to -8.5. In terms of individual locations, Site 747 basalts have high 87Sr/86Sr ratios (0.7056 to 0.7058), whereas  $\varepsilon_{Ne}$  ranges from 2.7 to -4, overlapping with the most enriched basalts from Kerguelen Island (Fig. 11). In comparison,  $\varepsilon_{Nd}$  is higher and more variable for basalts from Sites 749 ( $\epsilon_{Nd} = 1.9$  to 4.5) and 750 ( $\epsilon_{Nd}$ = 1.4 to 5.2). Site 750 tholeiites differ from those of Site 749 in having higher <sup>87</sup>Sr/<sup>86</sup>Sr ratios. Acid-leaching experiments (Fig. 11) and the high Sr contents of Site 750 tholeiites (Fig. 6) suggest that this is only partly caused by alteration. In marked contrast to the tholeiites from the central and northern part of the Kerguelen Plateau, which have Sr and Nd isotopic compositions within the mantle array, isotopic analysis of a tholeiite from Site 738 (Alibert, 1991), at the southern tip of the Kerguelen Plateau, gave extreme 87Sr/86Sr and  $\varepsilon_{Nd}$  values of 0.7090 and 8.5, respectively (Fig. 13), well outside the range of oceanic (hotspot-related) basalts.

Although the Pb isotopic compositions of basalts from Sites 748 and 749 overlap with data for Kerguelen Island, the tholeiites from Sites 738, 747, and 750, in comparison, are characterized by lower <sup>206</sup>Pb/<sup>204</sup>Pb (17.47 to 17.74), despite major differences in their Nd and Sr isotopic compositions. The single analysis of the Site 738 tholeiite (Alibert, 1991) has a very high <sup>207</sup>Pb/<sup>204</sup>Pb ratio (15.71), similar to values observed in continental lamproites from western Australia and leucitites from Gaussberg volcano, Antarctica (Fig. 12). Table 1. Major and trace element analyses of Leg 120 drill and Marion Dufresne (MD 48) dredge samples from the Kerguelen Plateau, Indian Ocean.

Hala	7470	7470	7470	7470	7470	7470	7470	7470	7470	7470	7490	7490	7490	7400	7400	7400	7400	7400
Core, section	11R-1	12R-2	12R-4	13R-3	14R-1	14R-2	15R-1	16R-2	16R-4	16R-5	79R-5	79R-6	82R-CC	12R-2	12R-3	12R-3	12R-4	15R-2
Interval (cm)	6-8	119-121	36	144-148	30-34	16-21	26-32	/8-80	50-53	103-106	137–139	90-94	0-2	64-69	82-84	117-120	138-140	118-123
Major elemen	ts:																	
SiO <sub>2</sub>	51.86	50.54	50.61	50.54	49.72	50.77	49.68	50.35	49.54	50.90	50.40	49.24	51.23	54.92	52.82	52.60	52.33	53.75
110 <sub>2</sub>	17.42	1.55	15.05	16.57	16.53	16.87	16.76	16.52	16 33	17.23	18.06	17.50	17.83	12 00	15.58	14.01	1.40	1.44
FeaOa	10.17	11 38	12.01	10.37	11 42	12.17	10.70	10.52	10.33	10.62	8 69	8 77	7.61	11.00	11 47	13.26	12 50	11 55
MnO	0.17	0.17	0.20	0.15	0.15	0.14	0.17	0.21	0.21	0.16	0.08	0.09	0.09	0.11	0.16	0.17	0.15	0.19
MgO	6.22	7.53	7.65	7.85	7.97	6.94	8.50	8.38	9.11	7.47	7.01	7.54	6.98	8.18	6.80	5.82	5.85	6.35
CaO	9.80	11.04	11.57	10.55	10.16	7.89	9.18	9.60	10.49	9.59	6.73	6.71	6.81	6.94	8.36	9.47	9.12	9.79
Na <sub>2</sub> O	2.73	2.43	2.31	2.59	2.50	2.76	2.01	1.92	2.13	1.88	4.61	4.72	4.80	3.27	3.31	2.92	2.77	2.95
K <sub>2</sub> Ō	0.673	0.529	0.238	0.511	0.588	1.110	1.208	1.778	0.893	1.998	0.497	0.557	1.318	2.211	0.628	1.179	1.004	0.472
P <sub>2</sub> O <sub>5</sub>	0.214	0.187	0.200	0.213	0.21	0.211	0.199	0.172	0.179	0.158	1.128	1.122	1.103	0.161	0.178	0.168	0.162	0.171
Total	101.29	101.00	101.47	101.49	100.93	100.83	99.85	100.93	100.44	101.22	100.02	98.99	100.49	100.89	100.85	101.00	99.46	101.14
LOI	1.88	1.26	0.67	1.46	2.05	3.00	5.20	3.00	2.73	3.63	5.66	7.14	4.86	2.60	1.14	1.72	1.13	0.13
CIPW norms:	1.02														0.77	0.54	0.07	
Qtz	1.83										0.41				0.67	0.54	2.27	3.47
Cor	2.08	3 13	1.41	3 02	3 17	6.56	7.14	10.51	5 28	11.01	2.04	3 20	7 70	12.07	2 71	6.07	5 02	2 70
Ab	23.10	20.56	19.55	21.92	21.16	23.36	17.01	16.25	18.02	15.01	39.01	30.04	40.62	27.67	28.01	24 71	23 44	24.06
An	33.30	30.21	30.00	32.09	32.15	30.37	33.15	31.22	32.37	32.68	26.02	24.93	23.22	14.24	25.81	21.65	23.27	24.88
Ne	00100	00121	20100	04107	Durio	00101	00110	01100	04001	0.000	LOIOL	21170	Lach Charles	11121	20101	21100	20121	24.00
Di	11.38	19.09	21.37	15.37	13.77	6.07	9.03	12.46	15.01	11.33		0.83	2.69	15.62	11.91	20.12	17.32	18.58
Hy	20.71	19.09	22.16	18.01	17.22	25.00	22.62	17.45	15.26	19.40	12.78	4.34	2.04	21.98	24.43	20.52	20.94	20.33
OI		2.59	0.30	4.64	6.51	2.08	5.22	7.51	8.97	4.66	8.66	15.58	14.43	2.98				
Mt	1.76	1.96	2.07	1.88	1.97	2.10	1.88	1.86	1.77	1.83	1.50	1.51	1.31	1.90	1.98	2.29	2.16	1.99
llm	3.86	2.94	3.10	3.13	3.19	3.74	2.39	2.36	2.45	2.30	5.34	5.20	5.17	2.11	2.92	2.66	2.66	2.73
Ар	0.50	0.45	0.46	0.49	0.49	0.49	0.40	0.40	0.41	0.37	2.61	2.60	2.50	0.37	0.41	0.39	0.38	0.40
XRF:	1010102-7	12127	12.02	12222	2025	1020125	101121	5050	2225	2327	100210	2218-24V	19275	127-22	1212	75714	12727	15.72
Nb	11.8	7.4	9.3	10.2	9.2	10.5	7.9	8.1	7.7	7.0	123.4	121.9	125.5	4.3	5.7	4.7	5.2	5.8
Zr	159.0	112.7	122.6	132.4	135.2	152.8	101.1	98.1	100.4	96.7	605.0	599.3	621.6	54.5	93.6	84.0	87.9	91.2
Y Sr	265.3	23.0	20.9	24.2	21.5	227.0	23.3	22.8	240.4	23.2	1226.0	1120.8	1125.2	137.6	30.1	202.4	30.2	31.3
Rh	8.4	11.2	243.9	92	67	18.9	17.5	8.1	240.4	18.5	7 3	7.8	7.8	18.2	12.8	41 1	212.0	9.9
Ga	22.9	20.0	19.1	21.3	19.7	22.7	17.2	18.4	18.4	22.0	18.2	18.7	19.4	16.1	22.7	20.9	22.2	23.9
Zn	130.4	92.1	93.4	97.9	93.2	98.6	79.2	77.9	71.1	65.9	76.6	79.0	86.5	73.2	121.1	103.4	97.7	97.4
Ni	55.6	69.1	65.5	75.0	73.8	50.9	107.9	100.8	107.0	71.6	224.3	181.7	210.7	52.2	29.7	30.6	30.0	31.2
v	377.3	306.1	299.3	250.0	265.4	326.2	194.6	187.7	211.3	201.0	179.0	169.5	168.1	199.3	278.8	234.0	247.3	243.5
Cr	182.2	278.6	224.0	82.6	90.5	114.2	351.2	371.1	359.6	401.4	274.2	166.0	218.1	235.0	274.3	237.8	235.4	229.2
Ba	142.9	160.7	147.7	250.3	246.7	228.8	187.5	205.5	175.9	229.3	1351.0	1661.0	1449.0	171.7	120.8	98.8	104.7	107.4
La	14.0	11.7	10.7	15.5	14.8	21.2	13.0	14.6	12.4	13.4	109.5	108.9	109.3	4.5	6.1	8.2	5.8	6.2
Ce Nd	30.0	24.5	30.3	32.9	37.8	48.2	33.3	30.3	30.5	31.3	214.9	218.3	220.3	6.2	10.4	14.0	14.8	19.3
ING	19.1	14.5	15.1	17.5	18.5	24.5	17.0	15.0	15.1	10.4	69.0	88.0	09.4	0.2	11.5	9.2	10.9	9.7
INAA:																		
La	13.2		12.5						13.3	12.3		105.0	106.0					
Ce Nd	32.0		28.0						28.0	25.5		102.0	230.0					
Sm	18.5		17.7						2.50	2 50		14.20	12.80					
Fu	1.38		1 29						1.07	1.03		3 31	3 58					
Th	0.67		0.78						0.63	0.65		1 25	1 38					
Yb	2.35		2.76						2.17	1.89		1.8	2.21					
Lu	0.30		0.37						0.31	0.28		0.26	0.26					
Та	0.66		0.54						0.56	0.46		9.08	7.50					
Hf	3.84		3.05						2.51	2.34		12.40	10.80					
Th	1.29		1.09						1.08	0.80		11.50	10.50					
W	6.1		7.6						6.0	5.3		16.4	15.7					
Sc	48.0		43.5						36.7	33.8		18.7	19.4					

Table 1 (continued).

Hole Core, section Interval (cm)	749C 15R-5 125-127	749C 15R-6 97–100	749C 16R-7 69-76	750B 16R-4 47-51	750B 14R-1 47-50	750B 15R-2 88-92	750B 16R-6 54-58	750B 17R-1 5-7	750B 17R-3 23-26	750B 17R-3 50-54	MD48-86 DR 02-2	MD48-86 DR 05-1	MD48-86 DR 05-2	MD48-86 DR 05-3	MD48-86 DR 06	MD48-86 DR 07	MD48-86 DR 08-4
$\begin{tabular}{ c c c c c } \hline Major element \\ SiO_2 \\ TiO_2 \\ Al_2O_3 \\ Fe_2O_3 \\ MnO \\ MgO \\ CaO \\ Na_2O \\ CaO \\ Na_2O \\ R_2O \\ P_2O_5 \\ Total \end{tabular}$	s: 52.73 1.52 15.18 11.95 0.17 7.22 8.31 3.42 0.543 0.177 101.22	50.55 1.75 14.90 13.41 0.19 6.82 9.96 3.21 0.228 0.189 101.21	49.98 0.92 19.01 9.15 0.13 7.01 12.11 2.48 0.081 0.086 100.96	51.04 0.73 15.44 11.96 0.19 9.33 11.00 1.97 0.092 0.065 101.82	42.65 0.93 17.75 26.12 0.06 6.43 0.34 0.45 5.283 0.087 100.10	50.38 0.71 15.42 11.95 0.19 10.15 9.97 1.81 0.109 0.055 100.74	50.33 0.76 16.01 11.17 0.18 8.47 11.36 2.05 0.095 0.061 100.49	50.40 0.75 15.71 11.68 0.18 8.99 11.31 1.99 0.053 0.069 101.13	49.21 1.17 15.77 14.62 0.18 8.36 9.01 1.99 0.189 0.113 100.61	49.75 1.23 16.55 13.96 0.13 9.87 6.55 2.09 0.259 0.108 100.50	50.74 2.47 17.92 9.71 0.17 5.72 6.65 3.62 2.332 0.447 99.78	50.02 1.72 16.66 11.79 0.19 4.99 9.41 3.32 1.083 0.332 99.51	50.67 1.66 15.84 11.75 0.19 6.18 9.82 3.44 0.569 0.211 100.33	50.61 1.61 16.00 10.26 0.14 5.80 9.58 3.26 1.020 0.228 98.51	49.05 2.38 17.07 14.14 0.20 3.88 8.76 3.67 0.895 0.531 100.58	50.08 1.70 16.39 12.44 0.23 5.94 9.57 3.49 0.632 0.208 100.68	49.74 2.26 15.13 13.85 0.21 4.91 9.56 3.07 1.440 0.292 100.46
LOI CIPW norms: Qtz Cor Or Ab An Ne	1.35 3.21 28.94 24.47	0.94 1.35 27.16 25.58	0.90 0.48 20.99 40.51	0.98 0.54 16.67 33.02	7.30 10.88 31.22 3.81 1.12	0.64 15.32 33.64	1.18 0.56 17.35 34.21	0.31 16.84 33.79	4.31 1.12 16.84 33.55	4.22 0.33 1.18 1.53 17.69 31.79	1.97 13.78 30.63 25.77	1.44 6.40 28.09 27.37	0.90 3.36 29.11 26.11	0.97 6.03 27.59 26.02	1.84 5.29 31.06 27.47	1.16 3.73 29.53 27.20	0.83 8.51 25.98 23.26
Di Hy Ol Mt Ilm Ap	12.80 24.47 0.93 2.06 2.89 0.41	18.65 15.02 6.20 2.31 3.32 0.44	15.50 14.09 5.07 1.58 1.75 0.20	17.25 28.30 1.40 2.06 1.39 0.15	27.05 17.26 4.51 1.77 0.20	12.59 32.41 1.57 2.06 1.35 0.13	17.77 24.63 1.49 1.93 1.44 0.14	17.86 25.25 2.46 2.02 1.42 0.16	8.58 32.56 1.69 2.52 2.22 0.26	41.76 2.41 2.34 0.25	3.51 6.73 11.11 1.68 4.69 1.04	14.23 9.11 7.22 2.04 3.27 0.77	17.49 9.89 7.67 2.03 3.15 0.49	16.44 11.47 4.72 1.77 3.06 0.53	10.53 7.44 9.37 2.44 4.52 1.23	15.64 6.62 11.01 2.15 3.23 0.48	18.57 7.43 8.15 2.39 4.29 0.68
XRF: Nb Zr Y Sr Rb Ga Zn Ni V Cr Ba La Ce Nd	5.8 91.1 29.3 213.5 12.3 22.7 115.1 29.5 271.3 259.5 113.8 6.5 15.2 10.6	$\begin{array}{c} 6.3 \\ 100.0 \\ 32.4 \\ 227.7 \\ 6.9 \\ 21.2 \\ 98.9 \\ 39.3 \\ 317.0 \\ 149.6 \\ 93.9 \\ 6.5 \\ 21.9 \\ 13.1 \end{array}$	$\begin{array}{c} 3.9\\ 41.2\\ 17.4\\ 225.0\\ 4.4\\ 18\\ 60.6\\ 69.6\\ 187.8\\ 278.4\\ 49.4\\ 4.1\\ 9.0\\ 5.5\end{array}$	26.9 19.1 119.6 5.4 18.5 79.0 103.0 234.7 100.0 47.2 5.2	30.0 13.6 40.8 22.0 17.4 76.3 156.0 278.4 131.8 192.3 6.1	24.2 18.3 112.9 4.9 18.0 79.9 108.5 220.5 92.7 39.8 3.5	26.7 19.6 128.4 4.5 17.7 82.7 115.6 241.4 100.6 47.1 5.4 6.8 2.1	26.3 20.6 122.5 2.0 19.1 80.8 112.9 238.9 106.9 34.5 2.6 3.9 3.6	4.2 46.9 24.6 192.5 9.0 19.8 93.4 120.3 269.4 192.5 30.4 6.2 8.0 7.0	3.4 46.5 21.3 129.3 6.6 21.2 84.0 108.7 290.3 241.4 20.9 4.1	$\begin{array}{c} 12.7\\ 199.8\\ 36.1\\ 368.2\\ 69.4\\ 26.8\\ 213.3\\ 27.6\\ 268.6\\ 168.6\\ 140.4\\ 17.9\\ 47.5\\ 27.5 \end{array}$	5.6 106.7 30.0 226.3 25.8 24.6 184.8 71.6 294.3 340.3 98.8 9.7 20.2 14.8	5.6 107.1 33.2 232.2 9.2 20.6 142.3 49.5 286.6 309.5 57.2 7.2 20.4 13.5	6.0 102.0 34.4 217.5 39.3 22.7 175.9 51.9 262.5 333.2 94.9 9.1 17.9 13.6	$11.3 \\ 179.4 \\ 47.1 \\ 342.4 \\ 14.0 \\ 28.5 \\ 222.5 \\ 51.8 \\ 348.5 \\ 186.1 \\ 137.6 \\ 15.4 \\ 34.7 \\ 23.5 \\ 186.1 \\ 137.6 \\ 15.4 \\ 34.7 \\ 23.5 \\ 186.1 \\ 10.4 $	$\begin{array}{c} 7.2\\ 107.3\\ 28.6\\ 239.2\\ 11.3\\ 23.9\\ 175.4\\ 54.7\\ 305.7\\ 315.3\\ 67.3\\ 9.1\\ 19.6\\ 13.7 \end{array}$	$\begin{array}{c} 12.5\\ 175.0\\ 42.6\\ 169.1\\ 28.0\\ 24.6\\ 191.6\\ 64.7\\ 378.5\\ 170.1\\ 180.2\\ 17.4\\ 42.5\\ 20.0 \end{array}$
INAA: La Ce Nd Sm Eu Tb Yb Lu Ta Hf Th W Sc	$\begin{array}{c} 6.8\\ 16.1\\ 11.9\\ 3.64\\ 1.24\\ 0.69\\ 2.71\\ 0.39\\ 0.40\\ 2.16\\ 0.67\\ 6.2\\ 34.5 \end{array}$		$\begin{array}{c} 3.2\\ 9.6\\ 6.3\\ 2.06\\ 0.83\\ 0.53\\ 1.50\\ 0.22\\ 0.15\\ 1.08\\ 0.24\\ 1.9\\ 30.8 \end{array}$			$\begin{array}{c} 2.5\\ 5.6\\ 4.5\\ 1.41\\ 0.59\\ 0.43\\ 2.04\\ 0.30\\ 0.10\\ 0.76\\ 0.21\\ 3.3\\ 40.9\end{array}$			4.0 8.9 6.3 2.32 0.84 0.58 2.57 0.34 0.19 1.16 0.49 4.0 38.9		$\begin{array}{c} 17.8 \\ 47.0 \\ 21.0 \\ 5.30 \\ 1.74 \\ 0.80 \\ 2.11 \\ 0.33 \\ 0.54 \\ 4.10 \\ 2.50 \\ 1.8 \\ 29.9 \end{array}$		$\begin{array}{c} 10.4\\ 21.0\\ 13.4\\ 4.70\\ 1.57\\ 0.81\\ 2.85\\ 0.47\\ 0.40\\ 2.40\\ 1.50\\ 3.9\\ 42.2 \end{array}$		$19.8 \\ 52.0 \\ 28.0 \\ 7.40 \\ 2.19 \\ 1.07 \\ 4.00 \\ 0.70 \\ 0.68 \\ 4.20 \\ 2.60$	$\begin{array}{c} 7.0 \\ 17.3 \\ 11.6 \\ 3.77 \\ 1.29 \\ 0.82 \\ 2.27 \\ 0.38 \\ 0.39 \\ 2.41 \\ 0.70 \\ 3.2 \\ 35.6 \end{array}$	$\begin{array}{c} 16.0\\ 36.0\\ 19.6\\ 5.60\\ 1.49\\ 1.10\\ 3.30\\ 0.55\\ 0.85\\ 4.09\\ 2.00\\ 2.2\\ 40.7 \end{array}$

LOWER CRETACEOUS VOLCANIC ROCKS

Note: LOI = loss on ignition.



Figure 5. Plots of the MgO vs. Ti/Zr ratio for samples from (A) Kerguelen Plateau, (B) Naturaliste Plateau and Bunbury Basalt, and (C) Rajmahal tholeiites.

On a Pb-Sr diagram (Fig. 13A), all of the Kerguelen Plateau samples lie above the Indian Ocean MORB field, with some falling in the Kerguelen Island field. In the Pb- $\varepsilon_{Nd}$  plot, a number of the plateau samples fall between the MORB and Kerguelen Island fields. The low 206Pb/<sup>204</sup>Pb tholeiites from Sites 738, 747, and 750 plot roughly between the fields of Indian Ocean MORB and Western Australian lamproites (Fig. 13).

#### **Rajmahal Basalts**

The Sr and Nd isotopic data for the Rajmahal basalts define two distinct groups, with  ${}^{87}$ Sr/ ${}^{86}$ Sr and  $\varepsilon_{Nd}$  values ranging from 0.7037 to 0.7053 and 5.1 to 2.9 (Group I) and 0.7064 to 0.7075

and -1.5 to -4.4 (Group II), respectively. This mirrors the variations in incompatible elements (Fig. 8). Most of the Group I tholeiites have Sr and Nd isotopic compositions similar to Site 749 basalts from the Kerguelen Plateau, although two Rajmahal samples are distinguished by higher 87Sr/<sup>86</sup>Sr values (Fig. 11). The Rajmahal basalts are characterized by low <sup>206</sup>Pb/<sup>204</sup>Pb (17.30 to 17.69) and high <sup>207</sup>Pb/<sup>204</sup>Pb (15.59 to 15.62) ratios (Fig. 12).

### DISCUSSION

Kerguelen Plateau basalts have Sr and Nd isotopic compositions lying outside the field for Indian Ocean MORB, thus excluding the possibility that the plateau represents a piece of old, uplifted normal ocean crust (Fig. 13). With the exception of Site 738, the Sr and Nd isotopic data are within the range of hotspot-related oceanic basalts, those from Sites 747 and 748 overlapping with the field for the most recent basalts from Kerguelen Island (Fig. 13). However, basalts from Kerguelen Island are characterized by higher 206Pb/204Pb ratios than plateau samples from Sites 738, 747, 750, and MD48 DR2 (Fig. 13). The low 206Pb/204Pb values of these particular samples may indicate the involvement of components derived from the continental lithosphere (discussed below). Despite the geochemical complications that have apparently arisen through interactions of the plateau source with continental lithosphere, a number of samples do show evidence for the involvement of Kerguelen Island-type mantle (Fig. 13), supporting the hotspot (Kerguelen-Heard plume [KHP]) hypothesis for most of the plateau, as do the plate reconstructions (Fig. 1, inset; Davies et al., 1989; Storey et al., 1989). One important difference between the younger basalts of Kerguelen and Heard islands, which were erupted in an intraplate environment (e.g., Storey et al., 1988), and the plateau samples is that the latter have lower abundances of such incompatible elements as Th, Nb, Ta, Zr, and LREE, implying much higher degrees of melting of the KHP source during formation of the Kerguelen Plateau. In this respect, the Kerguelen Plateau samples are similar to basalts from other large-volume, hotspot-related oceanic magmatic provinces, such as the Caribbean Plateau, the Ontong Java Plateau, and Nauru Basin (e.g., Floyd, 1989). These observations are consistent with a near- or on-ridge setting for the Kerguelen Plateau (perhaps analogous to Iceland), an environment that would favor extensive melting of an upwelling mantle plume (McKenzie and Bickle, 1988; Storey et al., 1991). It has also been suggested that the largest oceanic plateaus (e.g., Kerguelen Plateau, Ontong Java Plateau) may have formed by melting of the large heads thought to form during the initiation of mantle plumes (Richards et al., 1989).

Low 206Pb/204Pb and sometimes high 207Pb/204Pb ratios are a feature of old continental mantle lithosphere (e.g., Nelson et al., 1986; Hawkesworth et al., 1990). An extreme expression of this lithospheric signature is shown by western Australian lamproites (Figs. 12 and 13). One explanation for the low <sup>206</sup>Pb/<sup>204</sup>Pb values of some Kerguelen Plateau tholeiites is that Gondwana continental mantle lithosphere was incorporated into the Kerguelen plume source during the formation of the plateau. Contamination by lithospheric components was also suggested on the basis of the high La/Ta (Nb) and Th/Ta (Nb) ratios of some plateau tholeiites (Storey et al., 1989). The occurrence of Indian Ocean MORBs with low 206Pb/204Pb ratios has likewise been considered in terms of contamination by lithosphere detachment (e.g., Mahoney et al., 1989, in press). Thermal mobilization of the Gondwana mantle lithosphere by the KHP is plausible given the small size of the Indian Ocean basin (700 km) at the time of plateau formation and the 1000-2000-km-diameter heads attributed to mantle



Figure 6. MORB-normalized incompatible element plots and (insets) chondrite normalized REE plots for (A, B, C) Leg 120 basement tholeiites, (D) Leg 119 basement tholeiite, and (E, F) MD48 dredge samples. Leg 119 data from Alibert (1991). Normalizing values are from Sun and McDonough (1989).

plumes (White and McKenzie, 1989). Although plate reconstructions for Gondwana indicate that the Kerguelen Plateau must be essentially oceanic (Fig. 1; de Wit et al., 1988; Powell et al., 1988), it should be noted that the extreme isotopic compositions shown by a Site 738 tholeiite may indicate that the southernmost part of the plateau is underlain by continental crust (Alibert, 1991), possibly in a setting analogous to the North Atlantic Rockall Plateau (Morton and Taylor, 1987; Merriman et al., 1988) and Vøring Plateau sequences (Viereck et al., 1988).

Among the continental volcanism that accompanied the breakup of eastern Gondwana and the formation of the Kerguelen Plateau, the Rajmahal basalts are the most signifi-



Figure 6 (continued).

cant in terms of volume. The eruption of both may have been partly contemporaneous. Group I Rajmahal tholeiites with low La/Nb ratios share strong compositional similarities with some Kerguelen Plateau basalts (Fig. 14), suggesting a direct contribution to Raimahal volcanism by the KHP, in contrast to the conclusion reached by Mahoney et al. (1983). The Sr and Nd isotopic compositions of the Group I tholeiites supports this contention, although, their high 207Pb/204Pb ratios also indicate the presence of continental lithosphere components; the preferential enrichment of Pb, relative to Sr and Nd, in continental lithosphere makes Pb a sensitive indicator of such contamination. The range in La/Nb and K/Rb ratios of Group I tholeiites (Fig. 8) also suggests some contamination by continental lithosphere or melts derived therefrom. A probable example of the latter are the Lower Cretaceous Indian lamprophyres, which are characterized by very high La/Nb ratios and moderate to high K/Rb ratios (up to 1000;

Sample	EL55 12	EL5512-2	EL5512-7	EL5512-8	EL5512-9	EL5512-7*	EL5512-8*	BB1	BB2	BB3	BB4	BB5	BB6	BB7
Major ele	ements:			2304.022	100 - 100	10000	Sector C	POINT PRIM	1910/01/1		mious	-	messard	
SiO <sub>2</sub>	56.38	59.94	56.63	54.65	55.63	55.03	54.82	51.91	52.16	53.25	51.69	52.66	52.98	53.11
TiO <sub>2</sub>	1.81	1.80	1.80	1.88	1.85	1.69	1.74	1.67	2.05	2.05	1.77	1.96	1.31	1.31
$Al_2O_3$	17.13	15.04	16.90	18.02	17.59	16.82	17.47	16.20	14.55	14.43	15.57	15.21	14.72	14.71
Fe <sub>2</sub> O <sub>3</sub>	11.91	9.86	12.06	13.18	12.16	13.62	13.02	11.38	12.46	12.26	11.67	12.13	11.87	11.80
MnO	0.20	0.12	0.18	0.20	0.19	0.16	0.17	0.17	0.19	0.20	0.18	0.19	0.20	0.20
MgO	5.04	4.04	4.99	4.98	5.00	5.08	4.80	5.80	5.71	5.81	5.71	5.57	6.16	6.54
CaO	3.62	4.02	3.52	3.87	3.76	3.37	3.97	10.31	9.67	9.64	10.12	9.77	10.46	10.51
Na <sub>2</sub> O	2.71	3.25	2.67	2.75	2.73	2.88	2.90	3.28	3.12	3.07	3.09	3.22	2.94	2.83
K <sub>20</sub>	1.01	2.31	1.14	1.01	1.02	1.11	0.88	0.23	0.38	0.48	0.33	0.39	0.39	0.40
$P_2O_5$	0.22	0.59	0.22	0.24	0.21	0.24	0.23	0.18	0.22	0.22	0.18	0.21	0.13	0.13
Total	100.03	100.97	100.11	100.78	100.14	100.00	100.00	101.13	100.51	101.41	100.31	101.32	101.16	101.53
CIPW no	rms:													
Qtz	15.9	14.4	16.1	12.9	14.7			0.4	2.4	3.5	1.3	2.2	2.1	2.1
Cor	5.5	1.3	5.4	6.0	5.7									
Or	6.0	13.7	6.8	6.0	6.0			1.4	2.2	2.8	1.9	2.3	2.3	2.3
Ab	22.9	27.5	22.6	23.3	23.1			27.8	26.4	26.0	26.2	27.3	24.9	24.0
An	16.5	16.1	16.0	17.6	17.3			28.8	24.6	24.2	27.7	25.9	25.8	26.3
Di								17.5	18.1	18.3	17.6	17.5	20.8	20.6
Hy	26.1	20.7	26.2	27.6	26.3			18.8	19.2	19.1	18.8	18.8	19.4	20.4
Mt	2.1	1.7	2.1	23	2.1			2.0	22	2.1	2.0	2.1	2.1	2.0
Ilm	3.4	3.4	3.4	3.6	3.5			3.2	3.9	3.9	3.4	3.7	2.5	2.5
Ap	0.5	1.4	0.5	0.6	0.5			0.4	0.5	0.5	0.4	0.5	0.3	0.3
XRF:														
Nh	4 5	8.8	57	5.0	53	5.0	5.0		67	64	5.1	59	33	3.0
Zr	133.9	203.4	133.8	139.5	139.9	121.0	130.0		141.4	141	115.1	133.4	99.4	97.7
Y	25.7	26.6	26.5	28.4	24.9	34.0	26.0		39.2	39.9	35.2	37.3	31.3	32.2
Sr	135.7	131.6	131.3	140.7	138.0	129.0	150.0		234.6	229.4	231.2	247.3	160.3	162.6
Rb	37.5	61.2	36.5	37.8	37.6	23.0	20.0		4.3	14.6	5.6	5.7	11.3	8.9
Ga	24.0	23.5	24.5	26.2	25.0	20.0	20.0		25.1	24.5	24.1	25.3	20.8	19.5
Zn	271.4	272.4	274.4	306.4	281.3				117.4	118.2	104.6	103.0	88.2	83.5
Ni	122.7	111.0	139.0	127.9	125.3				37.5	37.8	69.0	41.1	50.1	58.3
INAA.														
La	12.1	28.6	13.3	12.4	99			6.6	10.2	98	72	10.0	9.2	8.5
Ce	31.0	59.0	41.0	36.0	30.0			18.0	25.6	30.0	21.4	26.3	25.1	18.4
Nd	14 7	29.9	20.6	14.9	11.7			10.0	20.0	50.0		40.0		10.1
Sm	4 33	63	4 17	4 76	3.84			3 77	5 28	5 17	4.15	4 98		3 53
Eu	1.27	1.99	1.29	1.51	1 50			1.26	1.45	1 39	1.42	1 58	1 47	1.55
Th	0.86	1 36	0.98	1 11	0.88			1.20	1.45	1.57	1.42	1.50	1.47	1.55
Yb	2 71	4 41	2 53	2.87	1.80			3.07	3 50	3 56	2.86	3 21	2.93	2.84
Lu	0.29	0.52	0.29	0.28	0.24			0.4	0.46	0.41	0.38	0.37	0.40	0.42
Ta	0.33	0.62	0.32	0.34	0.32			0.4	0.40	0.41	0.50	0.57	0.40	0.42
Hf	4.6	5.5	29	37	3.4			3.6	4 5	4.5	3.8	5.0	3.0	32
Th	1 40	6.40	1.62	1.87	1.80			0.69	1.60	1 20	1 20	1.50	1 30	1.50
w	19	67	33	4.6	1.00			0.09	2.2	4.0	3.2	1.5	27	2.2
Sc	25.9	21.8	24.7	27.5	27.2			30.5	33.4	33.4	30.7	31 3	41 4	40.8
	hadad	21.0	24.1	41.5	21.2			30.5	55.4	55.4	50.7	91.9	41.4	10.0

Table 2. Major and trace element analyses of samples from the Naturaliste Plateau and the bundury basait,	Australia.
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Notes: Samples from the Naturalistic Plateau were collected during the *Eltanin* (EL) cruise; locations are shown in Figure 3. Samples marked with an asterisk (\*) are data taken from Coleman et al. (1982).

Middlemost et al., 1988). These probably represent low degree melts of the continental mantle lithosphere, as suggested by the presence of mantle phlogopite and high incompatible element (Ni and Cr) abundances (Middlemost et al., 1988).

Compared with Group I tholeiites, Group II tholeiites have higher  ${}^{87}$ Sr/ ${}^{86}$ Sr ratios, show strong enrichment in highly incompatible elements, and generally have higher La/Nb ratios (Fig. 8). One possibility is that Group II tholeiites represent crustal-contaminated Group I tholeiites. Simple mixing calculations show that contamination by approximately 10%–20% from the upper crust mimics quite well the observed incompatible element abundances of the Group II tholeiites (Fig. 14). Although there may be alternative explanations for the geochemical features of the Group II tholeiites, indirect support for the crustal contamination hypothesis is that the majority of the Rajmahal tholeiites fall on or slightly below the one atmosphere olivine-plagioclase-clinopyroxene cotectic, implying that most have been affected by lowpressure, upper crustal fractionation (<1–2 kbar). The samples from Naturaliste Plateau and the Bunbury Basalt show varying degrees of enrichment in the large ion lithophile elements (LILE) over normal MORB. Although isotopic data were not available at the time of writing, these samples all appear to have strong continental lithosphere signatures, as suggested by their high La/Nb ratios (Fig. 7) This is not inconsistent with the suggestion of Heezen and Tharp (1973) that the Naturaliste Plateau may consist of stretched continental lithosphere, near the ocean-continent divide.

#### CONCLUSIONS

Plate reconstructions are consistent with a KHP origin for the Kerguelen Plateau (possibly in a tectonic environment analogous to Iceland at the present day) and associated magmatism along the continental margin of India, Australia, and Antarctica. Furthermore, the geochemistry of the Kerguelen Plateau supports the hotspot hypothesis. Some plateau samples are consistent with a KHP source, whereas others



Figure 7. MORB-normalized plots and (insets) chondrite-normalized REE plots for (A) *Eltanin* dredge samples from the Naturaliste Plateau and (B) Bunbury Basalt tholeiites.

have compositions that are consistent with mixing between MORB and KHP mantle.

Samples from the central and southern parts of the plateau that have anomalously low <sup>206</sup>Pb/<sup>204</sup>Pb and high La/Nb ratios suggest contamination of the KHP source by components derived from the Gondwana lithosphere. This lithospheric component has its extreme expression in a tholeiite from Site 738 and may indicate that the southern end of the plateau is underlain by continental crust.

Compositional similarities between some Rajmahal tholeiites (Group I) and Kerguelen Plateau samples suggest a direct contribution to Lower Cretaceous continental volcanism by the KHP. Other examples (Rajmahal Group II tholeiites, Naturaliste Plateau, and Bunbury Basalt) show evidence for the substantial involvement of components derived from the continental lithosphere, possibly by crustal contamination of KHP magmas. The Indian lamprophyres may represent smalldegree melts of continental mantle lithosphere, with the KHP providing a possible heat source for this activity.

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### LOWER CRETACEOUS VOLCANIC ROCKS

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# Table 3. Major and trace element analyses of Rajmahal basalts, India.

Sample	RB88/2	RB88/10	RB88/15	RB88/20	RB88/21	RB88/29	RB88/30	RB88/34	RB88/35	RB88/42	BH-1	BH-6	TT-3
Major elements	:												
SiO <sub>2</sub>	53.20	52.30	51.80	52.90	50.70	52.90	51.40	57.30	53.00	51.10	54.02	52.6	52.09
TiO <sub>2</sub>	1.70	1.77	2.10	1.44	1.76	1.66	1.78	1.01	1.82	1.78	1.75	1.65	1.71
Al <sub>2</sub> Õ <sub>3</sub>	14.90	15.00	15.50	15.70	14.60	14.70	15.10	15.00	14.90	15.00	14.67	14.7	15
Fe <sub>2</sub> O <sub>3</sub>	11.00	11.90	12.10	9.90	11.90	10.70	12.00	8.80	11.40	12.10	10.92	10.7	12.01
MnO	0.17	0.19	0.15	0.16	0.18	0.16	0.19	0.12	0.18	0.18	0.17	0.17	0.2
MgO	6.10	6.50	5.30	7.10	6.60	6.00	6.20	8.00	5.50	6.80	5.91	5.9	6.48
CaO	9.50	10.60	10.10	9.80	10.50	9.20	10.8	6.30	8.90	10.80	9.45	9.2	10.76
Na <sub>2</sub> O	2.30	2.60	3.00	2.50	2.30	2.10	2.60	2.90	2.60	2.50	2.79	2.30	2.83
K <sub>2</sub> Õ	0.940	0.270	0.250	0.560	0.190	0.950	0.190	0.250	1.170	0.150	1.177	1.120	0.150
P205	0.220	0.190	0.230	0.200	0.150	0.240	0.190	0.120	0.260	0.190	0.231	0.230	0.185
Total	99.90	101.41	100.49	100.28	98.92	98.57	100.56	99.74	99.53	100.52	101.09	98.49	101.42
CIPW norms:													
Qtz	5.64	3.02	2.89	3.76	3.27	7.05	2.50	10.31	4.68	1.92	3.90	5.36	1.81
Or	5.56	1.60	1.48	3.31	1.12	5.61	1.12	1.48	6.91	0.89	6.96	6.62	0.89
Ab	19.46	22.00	25.39	21.16	19.46	17.77	22.00	24.54	22.00	21.16	23.61	19.46	23.95
An	27.56	28.47	28.10	29.97	28.96	27.89	28.98	27.18	25.54	29.27	24.04	26.49	27.79
Di	14.90	18.82	16.97	14.14	18.21	13.30	19.25	2.67	13.97	18.96	17.55	14.51	20.07
Hy	20.32	20.52	18.05	22.15	21.07	20.50	19.66	29.14	19.62	21.45	18.35	19.68	20.12
Mt	1.90	2.05	2.09	1.71	2.05	1.85	2.07	1.52	1.97	2.09	1.88	1.85	2.07
Ilm	3.23	3.36	3.99	2.73	3.34	3.15	3.38	1.92	3.46	3.38	3.32	3.13	3.25
Ap	0.51	0.44	0.53	0.46	0.35	0.56	0.44	0.28	0.60	0.44	0.54	0.53	0.43
XRF:													
Nb	8.7		6.4	5.7	6.5	7.6	5.7		10.2	4.6	8.4	9.9	
Zr	160.6	107.4	120.8	118.2	106.8	160.3	99.8	49.9	170.4	93.9	159.7	160.3	108.3
Y	35.3	30.6	30.9	25.7	31.5	33.8	30.6	26.7	36.8	29.6	34.8	33.9	31.0
Sr	332.2	249.1	264.4	328.7	245.2	322.8	227.1	148.2	324.3	223.3	325.3	321.4	237.9
Rb	22.6	7.2	2.0	17.3	1.9	31.7	6.5	5.9	32.8	4.1	23.1	22.1	0.3
Ga	22.6	22.8	24.8	22.5	22.5	21.7	22.8	22.6	21.8	24.0	21.6	21.1	22.2
Zn	101.5	106.7	107.6	91.6	100.9	103.6	103.3	138.4	105.5	104.7	94.6	105.0	93.7
Ni	31.6	69.4	51.7	37.4	63.6	34.1	59.2	147.4	29.6	65.4	17.7	29.5	68.6
v	222.3		324.0	208.6	271.3	230.5	265.3	130.0	223.2	271.8		229.4	
Cr	155.4		105.0	295.2	229.1	155.6	207.9	672.5	119.4	415.0		187.1	
Ba	316.4		122.7	201.5	79.3	306.5	55.6	63.5	371.9	44.4		308.6	
La	18.6		7.6	10.8	9.0	18.3	7.0	5.6	19.2	6.5		15.3	
Ce	44.6		20.5	31.7	18.2	43.2	15.5	8.0	41.7	14.0		38.1	
Nd	23.4		15.8	17.3	12.3	22.2	11.1	5.5	25.3	13.6		21.3	
INAA:													
La	16.4	7.9	8.8	11.9	7.4		7	3.9					7.5
Ce	38.2	21.1	22.8	27.7			18.9	8.3					21.6
Nd	21.0	13.7	13.1	13.8	11.7		1017						14.1
Sm	4.69	3.87	4.11	3.72	3.80		3.75	2.52					3.98
Eu	1.50	1.39	1.37	1.35	1.50		1.24	1.02					1.34
Tb	0.68	0.86	0.72	0.56	0.74		0.76	0.70					0.80
Yb	2.66	2.86	2.80	2.24	2.99		2.67	1.94					3.12
Lu	0.44	0.38	0.41	0.34	0.37		0.39	0.27					0.42
Та	0.55	0.46	0.46	0.40	0.38		0.34	0.15					0.47
Hf	4.40	2.97	3,50	3.03	3.40		2.9	1.48					3.02
Th	2.70	0.71	0.71	2.20	0.45		0.53	0.49					0.92
W	3.1	1.8	4.1	2.3	3.8		4.9	2.6					4.1
Sc	30.8	35.7	40.5	28.8	34.2		37.0	14.5					37.9
00000		5.531 N		1777-1780 (M	0.000.000		77.1.1.1.1.1.						24001240

Note: See Figure 4 for locations.

Table 3 (continued).

Sample	LH-1	GP-1	BL-3	KP-4	KP-6	KP-7	NP-1	NP-4	NP-7	NP-9	LA-2	LA-6
Major elements:				141 - 14 - 15 - 1		Sec. 1					1 and a lot	
SiO <sub>2</sub>	50.7	54.1		51.4	52.1	51.3	51.5		52.30	51.80	52.60	49.40
TiO <sub>2</sub>	1.55	1.74		2.21	1.49	2.22	1.76		1.82	1.69	2.09	2.18
Al <sub>2</sub> Õ <sub>3</sub>	15.2	14.7		14.6	15.6	14.7	14.4		15.00	14.40	13.70	14.00
Fe <sub>2</sub> O <sub>3</sub>	11.3	10.9		13.4	11.6	13.1	12.4		11.70	11.60	12.30	12.80
MnO	0.2	0.17		0.2	0.2	0.2	0.2		0.20	0.19	0.18	0.20
MgO	6	5.9		5.9	6.2	5.8	6.4		6.60	6.40	5.70	5.90
CaO	11.1	9.4		10.7	11.8	10.8	11		10.40	10.30	10.00	10.60
Na <sub>2</sub> O	2.50	2.30		2.5	2.5	2.6	2.5		2.50	2.50	2.30	2.40
K <sub>2</sub> Õ	0.170	1.100		0.230	0.190	0.250	0.210		0.270	0.310	0.370	0.290
P2O5	0.160	0.220		0.210	0.150	0.220	0.180		0.190	0.190	0.230	0.250
Total	98.93	100.50		101.29	101.86	101.14	100.55		101.04	99.49	99.49	98.14
CIPW norms:												
Qtz	2.47	6.64		3.18	2.28	2.72	2.63		3.72	3.83	7.12	2.34
Or	1.00	6.50		1.36	1.12	1.48	1.24		1.60	1.83	2.19	1.71
Ab	21.16	19.46		21.16	21.16	22.00	21.16		21.16	21.16	19.46	20.31
An	29.76	26.55		27.94	30.79	27.71	27.46		28.92	27.16	25.97	26.58
Di	19.99	15.34		19.68	22.06	20.22	21.38		17.61	18.67	18.29	20.14
Hy	18.25	19.40		19.87	18.23	18.94	19.7		21.04	20.08	18.75	18.89
Mt	1.95	1.88		2.31	2.00	2.26	2.14		2.02	2.00	2.12	2.21
Ilm	2.94	3.30		4.20	2.83	4.22	3.34		3.46	3.21	3.97	4.14
Ap	0.37	0.51		0.49	0.35	0.51	0.42		0.44	0.44	0.53	0.58
XRF:												
Nb	3.9	9.3	6.3	6.5	4.5	7.5	5.5	7.0	4.9	6.7	6.7	
Zr	95.2	159.1	84.8	127.0	83.8	131.3	101.6	118.1	102.5	107.3	114.5	
Y	29.5	33.6	26.8	35.0	26.6	35.8	31.7	36.4	30.6	30.5	33.2	
Sr	233.6	347.4	231.8	236.3	232.7	240.7	230.2	304.4	233.0	254.3	219.3	
Rb	4.0	30.2	3.9	4.2	5.0	6.1	2.9	8.1	2.5	6.5	5.9	
Ga	22.3	23.0	21.2	22.5	21.1	25.3	21.9	22.0	23.4	22.8	19.2	
Zn	95	101.7	91.8	110.2	88.9	111.3	101.9	114.2	105.8	97.7	102.6	
Ni	76.8	20.3	66.6	67.8	68.6	82.3	54.8	60.3	58.5	53.5	62.8	
V	277.0	235.8	274.4	294.4	267.1	305.5	287.9	328.2	265.9	255.1	304.0	
Cr	206.3	172.2	241.6	128.4	220.5	132.9	167.3	148.6	237.4	230.9	165.3	
Ba	82.1	265.3	97.6	89.3	92.4	102.8	143.4	246.0	74.6	83.0	80.3	
La	5.1	16.6	7.2	10.3	5.0	8.7	5.8	11.4	7.3	7.0	5.4	
Ce	17.5	39.1	11.8	26.1	19.0	23.9	21.1	23.7	22.2	22.2	18.2	
Nd	10.2	21.8	9.5	14.5	10.4	16.8	12.9	16.9	11.9	11.3	16.8	
INAA:												
La										7.8		
Ce										22.0		
Nd										13.5		
Sm										4.06		
Eu										1.39		
Tb										0.80		
Yb										3.26		
Lu										0.40		
Ta										0.46		
Hf										2.70		
Th										0.75		
W										3.8		
Sc							-			35.7		



Figure 8. The variation of the Zr/Y ratio in Rajmahal samples vs. (A) Ti/Zr, (B)  $K_2O$  %, (C) Rb (ppm), (D) K/Rb, and (E) La/Nb.



Figure 9. MORB-normalized plots and chondrite-normalized (inset) REE plots for (A) Group I Rajmahal tholeiites, (B) Group II Rajmahal tholeiites, and (C) Lower Cretaceous Indian lamprophyres; data from Paul and Potts (1981).

Table 4. Major and trace element analyses of alnoites from the Prince Charles Mountains, Antarctica.

Sample	D15	D18		
Major elements:				
SiO <sub>2</sub>	37.14	37.35		
TiO <sub>2</sub>	2.23	2.42		
Al <sub>2</sub> Õ <sub>3</sub>	9.47	10.55		
Fe <sub>2</sub> O <sub>3</sub>	12.11	12.54		
MnO	0.20	0.22		
MgO	16.11	13.88		
CaO	16.36	15.40		
Na <sub>2</sub> O	2.08	2.46		
K <sub>2</sub> Ô	3.29	3.51		
PaOr	0.88	1 30		
Total	99.87	99.63		
CIPW norms:				
An	6.8	7.4		
Lu	15.3	16.3		
Ne	9.5	11.3		
Di	(6,55)	1.8		
Cs	22.7	18.5		
01	39.1	34.0		
Mt	21	2.2		
Ilm	4.7	4.6		
Ap	2.0	3.0		
XRF:				
Nb	118.0	165.1		
Zr	235.7	437.7		
Y	37.5	48.7		
Sr	1003.9	1115.3		
Rb	129.9	148.0		
Th	10.6	11.7		
Zn	85.2	83.5		
Ni	281.2	238.9		
INAA:				
La	89.0	96.0		
Ce	223.0	225.0		
Nd	99.0	106.0		
Sm	10.2	11.0		
Eu	2.99	3 12		
Th	1.58	1.85		
Yh	2 35	2.29		
Lu	0.21	0.20		
To	0.21	0.30		
LIF	6.1	9.0		
Th	14.2	9.2		
111	14.5	12.5		
Se	21.0	13.5		
30	24.0	30.2		

Note: See Figure 1 for location.



Figure 10. MORB-normalized plots and chondrite-normalized (inset) REE plots for Prince Charles Mountains alnoites. Note the strong similarity with the Albian alkali basalt from Site 748.



Figure 11. <sup>87</sup>Sr/<sup>86</sup>Sr vs.  $\varepsilon_{Nd}$  for Kerguelen Plateau (data from Alibert, in press; Salters et al., this volume; M. Storey, unpubl. data) and Rajmahal basalts (data from Mahoney et al., 1983; M. Storey, unpubl. data). Data corrected for T ~ 110 m.y., except for Site 748, where T ~ 90 m.y. Site 750 sample acid leached in 6M HCl. Stippled area represents most enriched basalts from Kerguelen Island (Storey et al., 1988, and references cited therein). Data for normal Southeast Indian Ridge (SEIR) MORB from Dosso et al. (1988), Hamelin et al. (1986), and Michard et al. (1986).

### LOWER CRETACEOUS VOLCANIC ROCKS



Figure 12.  $^{206}$ Pb/ $^{204}$ Pb vs. (A)  $^{207}$ Pb/ $^{204}$ Pb and (B)  $^{208}$ Pb/ $^{204}$ Pb for Kerguelen Plateau and Rajmahal Trap tholeiites (data from Weis et al., 1989; Alibert, 1991; Salters et al., this volume; M. Storey, unpubl. data). Lines joining points represent the age correction for samples on which U and Pb were determined by isotope dilution. Fields for western Australian lamproites and Gaussberg leucitites from Nelson et al. (1986); field for Kerguelen Island basalts field from Dosso et al. (1979), Storey et al. (1988), and Gautier et al. (1990). NHRL = Northern Hemisphere Reference Line (Hart, 1984).



Figure 13. <sup>206</sup>Pb/<sup>204</sup>Pb vs. (A) <sup>87</sup>Sr/<sup>86</sup>Sr and (B)  $\varepsilon_{Nd}$  for Kerguelen Plateau tholeiites. Sources of data are the same as for Figures 11 and 12.



Figure 14. Comparison between MORB-normalized incompatible element plots for Group I Rajmahal tholeiite (RB 88/21) and Site 749 Kerguelen Plateau tholeiite. The addition of 20% of material with upper-crust composition (from Taylor and McLennan, 1985) to Group I Rajmahal tholeiite mimics incompatible element abundances of the Group II Rajmahal tholeiite illustrated (RB 88/2).