18. GEOCHEMISTRY AND PETROGRAPHY OF CRETACEOUS SILLS AND LAVA FLOWS, SITES 800 AND 802¹

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ABSTRACT

On the basis of their respective eruptive environments and chemical characteristics, alkalic dolerite sills from the northern Pigafetta Basin (Site 800) and tholeiitic pillow lavas from the Mariana Basin (Site 802) sampled during Ocean Drilling Program Leg 129 are considered to represent examples of the widespread mid-Cretaceous volcanic event in the western Pacific. Both groups of basic rocks feature mild, low-grade, anoxic smectite-celadonite-carbonate-pyrite alteration; late-stage oxidation is very limited in extent, with the exception of the uppermost sill unit at Site 800. The aphyric and nonvesicular Site 800 alkalic dolerite sills are all well-evolved mineralogically and chemically, being mainly of hawaite composition, and are similar to ocean island basalts. They are characterized by high contents of incompatible elements (for example, 300–400 ppm Zr), well-fractionated rare earth element patterns ([La/Yb]_N 18–21) and HIMU isotopic characters. They probably represent deep-sea, lateral, intrusive off-shoots from nearby seamounts of similar age. The olivine-plagioclase \pm clinopyroxene phyric tholeiitic pillow lavas and thin flows of Site 802 are nonvesicular and quench-textured throughout. Relative to normal-type mid-ocean ridge basalt, they are enriched in large-ion-lithophile elements, exhibit flat (unfractionated) rare earth element patterns and have distinctive (lower) Zr/Nb, Zr/Ta, La/Ta, and Hf/Th ratios. Overall they are compositionally and isotopically similar to the mid-Cretaceous tholeiites of the Nauru basin and the Ontong-Java and Manihiki plateaus. The Site 802 tholeiites differ from the thickened crustal segments of the oceanic plateaus, however, in apparently representing only a thin veneer over the local basement in an off-axis environment.

INTRODUCTION

One of the major problems encountered by previous attempts to sample Jurassic ocean crust in the western Pacific Ocean were the products of extensive mid-Cretaceous volcanism that filled many ocean basins with seamount-derived volcaniclastic debris and proved difficult to penetrate completely (Lancelot, Larson, et al., 1990). Of the three sites drilled during Ocean Drilling Program (ODP) Leg 129 (Fig. 1), only Site 801 penetrated Jurassic ocean crust; although both Sites 800 and 802 drilled acoustic "basement," these proved to be Cretaceous in age. The clear extrusive nature of the Cretaceous-aged pillow basalts at Site 802 was intriguing, especially as the magnetic anomaly patterns and water depth indicated that the ocean crust in this part of the western Pacific Ocean should be of Jurassic age (Lancelot, Larson, et al., 1990). Was this really Cretaceous ocean crust or another expression of mid-Cretaceous volcanism which spread a thin veneer of extrusives over Jurassic crust lying immediately below? Currently, we prefer the latter tectonic situation. There is little doubt that the mid-Cretaceous volcanic event was a major feature of the central-western Pacific (Larson, 1991) and is represented by different volcanic forms, such as reef-capped seamounts and ocean islands (with their attendent slumped volcaniclastic aprons), enormous oceanic plateaus, such as the Ontong-Java Plateau, and deep-sea sill-flow complexes, as in the Nauru Basin (for example, Winterer, 1973; Schlanger et al., 1981; Larson and Schlanger, 1981; Haggerty et al., 1982; Rea and Vallier, 1983; Schlanger and Moberly, 1986; Floyd, 1989). Leg 129 also demonstrated the existence of other deep-sea expressions of the mid-Cretaceous event in the form of sills intruding Early Cretaceous sediments (Site 800) and pre-Aptian lava flows (Site 802). This paper describes the petrography and geochemistry of these Cretaceous volcanics and briefly discusses their origin and tectonic setting. Isotopic and radiometric age data are discussed in companion papers by Castillo et al. (this volume) and Pringle (this volume), respectively.

LITHOSTRATIGRAPHY AND PETROGRAPHY

The recovery of igneous rocks from Sites 800 and 802 was limited, with only 7 m of dolerites (from 58 m penetration) from the former and 16.6 m of basalts (from 51 m penetration) from the latter. Although a number of volcanic cooling units were identified at each site (Lancelot, Larson, et al., 1990), in both cases the rocks sampled show little major lithological and chemical downhole variation, unlike the longer Jurassic crustal sequence at Site 801 (Floyd and Castillo, this volume). However, each location is distinctive in that Site 800 acoustic "basement" is composed of alkalic dolerite sills, whereas Site 802 exhibits tholeiitic basalt lavas. Various petrographic features displayed by the dolerites and basalts from these sites are shown in Figure 2.

Site 800 Alkalic Dolerite Sills

Mode of Emplacement and Age

Although actual contacts were not recovered, it is inferred that the three cooling units recognized at this site probably represent intrusive sills rather than lava flows (Lancelot, Larson, et al., 1990). However, interpretation of the emplacement mode of thick, massive, relatively coarse-grained cooling units only seen in incompletely cored sections can often be equivocal, such that intrusives can be confused with thick flows and visa versa. Nevertheless, in areas of rapid sedimentation, where magmas may only occasionally reach the seafloor as lava flows, sill-like intrusives are relatively common. For example, in the Guaymas Basin (Gulf of California) numerous intrusive bodies were emplaced into, and clearly hydrothermally altered, poorly consolidated claystones (Einsale et al., 1980). However, massive bodies of similar thickness and textural variety sampled at the mouth of the Gulf of California were all interpreted as flows rather than sills on various criteria, including absence of intrusive relationships (Griffin et al., 1983). Because of the lack of observable contacts in Hole 800A, the intrusive nature of the three dolerite units was inferred from the following features: (1) basal and interbedded cherty and radiolarite sediments were baked, bleached, and brecciated relative to lithologically similar, but unaltered, sediments higher in the sedimentary sequence; (2) the dolerites were medium textured throughout, with

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Figure 1. Map of the western Pacific Ocean showing the position of Leg 129 Sites 800, 801, and 802 (Lancelot, Larson, et al., 1990). Unshaded area represents normal Pacific ocean crust with magnetic lineation contours, whereas shaded areas represent thickened crustal sections and younger crust to the west of the Pacific subduction zones. Feature abbreviations are as follows: Caroline Islands (CI), Ontong-Java Plateau (OJP), Marshall Islands (MI), Nauru Basin (NB), Mid-Pacific Mountains (MPM), Shatsky Rise (SR), Hawaiian Ridge (HR) and Emperor Seamounts (ES).

finer-grained chilled margins, holocrystalline and nonvesicular quench textures and glassy margins characteristic of rapid cooling by seawater were conspicuously absent towards the margins; (3) the variable presence of late-stage secondary biotite and actinolite in all sills is interpreted as the result of contact metamorphism caused by sill intrusion. This feature is particularly well developed in the top of Unit 3 where both minerals replace earlier smectite-dominated alteration; and (4) magnetic results indicate that the least altered dolerites have a (primary) reversed polarity, whereas those portions affected by biotiteactinolite alteration exhibit a superimposed normal polarity which probably developed after initial emplacement.

The sills were emplaced into sediments of lowermost Cretaceous age (Berriasian) and on the basis of magnetic polarity correlations may have been intruded during pre- or early Aptian times (Lancelot, Larson, et al., 1990), whereas Ar-Ar dating (Pringle, this volume) indicate an intrusion age of 126 Ma. These data confirm the intrusive nature of the dolerites relative to the older enclosing sediments and their association with the mid-Cretaceous volcanic event.

Primary and Secondary Petrography

All three cooling units are represented by aphyric, nonvesicular, holocrystalline gray-green dolerites with medium-grained intersertal to interlocking textures. Background secondary alteration is dominated by smectite-carbonate-celadonite assemblages throughout. Directly below the basal sediments, however, Unit 1 is represented by a fine-grained highly altered basalt with a granular texture (chilled margin facies) and a reddish oxidized top that grades downwards into the more normal coarser-grained greenish dolerites.

Pale pink, poorly pleochroic titanaugite is the main mafic phase and forms stubby to elongate prims variable replaced by greenish smectites

and carbonate, leaving large skeletal, optically-continuous relicts. No olivine or its pseudomorphs appear to be present, although a few rare ragged subhedral crystals of a pale brown primary amphibole are present in Unit 2. Albite-twinned plagioclase laths commonly exhibit normal zoning (An55-60 to An30-35), although cores are often partially replaced by brownish smectite and opaque granular material (analcite or clays). Titanomagnetite is common throughout, forming small opaque granules in the finer grained portions of the dolerites and larger, variably corroded, subhedral grains and needles in the coarser interiors. At the top of Unit 1 this phase is completely oxidized to either red hematite or a limonitic stain that also impregnates adjacent smectites. Acicular apatite prisms are a common accessory throughout all units and the only primary phase to remain unaltered. Apart from patchy carbonate replacement, small alteration domains in the matrix are represented by either (1) granular dark brown smectite with pale green fibrous smectite, or (2) pale green fibrous smectite often with strong blue-green celadonite plates. In both cases, secondary biotite and actinolite may be seen concentrated in these domains and grow across them unaltered. This suggests that biotite and actinolite growth was later than the main smectite-dominated alteration and probably represents a higher grade of metamorphism superimposed on the dolerites by the thermal effects of adjacent intrusions. Biotite is more abundant than actinolite and may also nucleate on titanomagnetite grains. Rarely, biotite may be present as zoned euhedral plates with slightly darker rims relative to the optically uniform cores. In some highly altered domains, biotite plates may have ragged or splayed ends parting along the cleavage and are then partially replaced by blue-green celadonite or green smectite. This may reflect a subsequent alteration during cooling after thermal metamorphism.

The apparent lack of olivine and the generally intermediate nature of the plagioclase (andesine-labradorite) indicate that the dolerites are relatively evolved and (as substantiated by their major element chemistry) have an overall hawaiite composition rather than that of an alkali basalt.

Site 802 Tholeiitic Lava Flows

Mode of Emplacement and Age

The basalts of Hole 802A are clearly extrusive submarine flows being characterized by numerous thin, nonvesicular cooling units (about 1 m or less) with quench textures throughout and glassy spherulitic margins. The preservation of typical rapid cooling textural zones and, in a number of cases, curved glassy selvages (Fig. 44 in Lancelot, Larson, et al., 1990) indicates that the sequence here is predominantly composed of a series of pillow lavas. Hole 802A, however, terminated in a coarser-grained, hypocrystalline unit which is considered to be a thin (<5 m) flow and thicker than most of the measurable units above. The top of the lava sequence is not extensively reddened and indicates that the time interval, before being covered by sediment, was relatively short so that alteration by oxygenated seawater was restricted. Glassy fragments spilled off the uppermost pillow lavas during cooling by seawater are incorporated into the basal sediments (volcanogenic claystones) of late Aptian age, and provide a minimum age for the lavas. Ar-Ar dating of the basalts gives an age of about 115 Ma (Pringle, this volume) that corresponds to the uppermost Aptian (time scale by Harland et al., 1982).

Primary and Secondary Petrography

One of the features of the lava flow sequence at Hole 802A is that all cooling units are petrographically very similar, largely being composed of sparsely olivine-plagioclase or olivine-plagioclase-clinopyroxene microphyric tholeiitic basalts. However, up to about 20% of collected core is broadly aphyric, although even here scattered olivine and plagioclase phenocrysts may be encountered. Microphenocrysts are not uniformly distributed between or within cooling units, with some showing a concentration in the marginal highly quenched zones relative to the phenocryst-poor interiors. Larger







0.1 mm







0.1 mm

E



0.5 mm



0.1 mm



0.1 mm



Figure 2. Photomicrographs illustrating petrographic and textural features of Site 800 alkalic dolerite sills (numbers 1-6) and Site 802 tholeiitic flows (numbers 7-12). (ppl = plane polarized light; cp = cross polars). A. Small brownish glass inclusions in titanaugite prism, Sample 129-800A-60R-1, 43-44 cm (ppl). B. Carlsbad twinned and zoned plagioclase lath with core replaced by dark granular smectite and analcite, Sample 129-800A-57R-2, 89-90 cm (cp). C. Acicular apatite crystals, Sample 129-800A-60R-1, 71-74 cm (ppl). D. Large plate of poorly zoned actinolite with later fibrous fringe in brown smectite alteration domain, Sample 129-800A-58R-1, 46-47 cm (ppl). E. Primatic, fibrous actinolite growing unaltered across a brown smectite domain, Sample 129-800A-58R-1, 46-47 cm (ppl). F. Group of biotite plates in smectite alteration domain, with frayed terminations partially replaced by blue-green celadonite, Sample 129-800A-60R-1, 71-74 cm (ppl). G. Dark brown, partly coalesced spherulites nucleated on plagioclase microlites set in fresh glass (sideromelane), Sample 129-802A-58R-2, 145-149 cm (ppl). H. Smectite pseudomorphed olivine microphenocrysts in spherulitic zone of lava, Sample 129-802A-58R-2, 145-149 cm (ppl). I. Double-fan variolite growth on plagioclase microlite surrounded by opaque magnetite granules, Sample 129-802A-58R-4, 55-57 cm (ppl). J. Coarse sheaf and plumose variolites of clinopyroxene typical of the interior of flows, Sample 129-802A-57R-2, 126-128 cm (ppl). K. Group of clinopyroxene and zoned plagioclase phenocrysts, sample 129-802A-57R-2, 126-128 cm (cp). L. Glomerophyric cluster of anhedral clinopyroxene grains and plagioclase laths in quenched matrix, Sample 129-802A-57R-2, 126-128 cm (cp).

dimensioned plagioclase phenocrysts (about 1 mm in length) may also be observed and are termed "megaphenocrysts" to distinguish them from the common smaller type.

Textural variation within units is largely the result of differential quenching by seawater, together with the random distribution of glomerophyric clumps and individual phenocrysts. Rapidly chilled selvages are glassy with variably coalesced spherulites that may often nucleate on plagioclase microlites, whereas the matrix of flow interiors grade from sheaf and plumous variolites to coarser intersertal and granular textures that may still exhibit quench morphologies (Fig. 2). Microphenocrysts of olivine are always replaced by brown fibrous smectite and occur as individual crystals scattered throughout both glassy and quenched crystal matrices. Plagioclase is often fresh and may be present as either broken and magmatically corroded zoned megacrysts (with cores of around An₈₀), or, more usually, as microphenocrysts (commonly An₆₅₋₇₀) in glomerophyric clumps with clinopyroxene. Clinopyroxene (augite) is rarely seen as individual "free" microphenocysts being invariably concentrated in glomerocrystic clumps. There are textural phenocryst associations that often characterize the more crystallized variolitic zone just below the glassy selvage of most units. On the one hand, there may be scattered, subhedral plagioclase megaphenocrysts and relatively abundant olivine microphenocrysts, whereas on the other are finer-grained glomerophyric clumps of quenched plagioclase laths and granular clinopyroxene. Both are set in a poorly birefringent matrix of brown plumous variolites and scattered magnetite granules. The two sizes of phenocryst imply different cooling histories, with the plagioclase megaphenocrysts and olivine developing in the magma chamber and the glomerophyric clumps at a higher level possibly just prior to actual extrusion. Any remaining melt on contact with seawater will then rapidly quench to produce a variolitic and/or glassy matrix. The textural diversity and association of glomerophyric clumps and variolites is common to all flows and is only replaced in the coarser interiors of larger cooling units with intersertal textures composed of granular clinopyroxene and serrated plagioclase laths, such as the thick olivine-plagioclase microphyric flow at the base of Hole 802A.

In terms of secondary alteration, the Cretaceous flows are remarkably fresh, being characterized by mainly smectite-carbonate-pyrite assemblages, both within the bulk rock and veins. This assemblage is typical of low-grade anoxic conditions. Color changes in various shades of gray within the core represent alteration haloes usually associated with areas of veining and more extensive matrix alteration by smectites. The preservation of fresh yellow sideromelane in some pillow selvages (for example, Sample 129-802A-58R-2, 99–102 cm) is rare, as most glass is replaced by brownish smectite. Olivine microphenocrysts are always pseudomorphed by brown smectites or in some glassy selvages by green smectite. Carbonate replacement is later than smectite alteration and occurs in a random, patchy fashion through the matrix. Plagioclase is sometimes replaced by carbonate, whereas clinopyroxene always remains fresh.

GEOCHEMISTRY

Major oxide and trace element data for Hole 800A dolerite sills and Hole 802A basalt lavas are shown in Tables 1 and 2, respectively. Analytical procedures and analysts are referenced in Floyd and Castillo (this volume). Geochemical data for both holes are plotted on the same diagrams to illustrate their main features and highlight the significant differences between the alkalic dolerites and tholeiitic basalts.

Apart from the main chemical differences between the alkalic dolerites and the tholeiitic basalts that are typical of their respective magma types, both groups feature very little chemical variation throughout the sampled sequences. This uniformity is also mirrored in the similarity of petrographic composition within each hole and may, in part, reflect the relatively short segment sampled in each case. However, small, relatively insignificant differences with depth (meters below sea floor, or mbsf) have been noted. The uppermost part of Hole 800A (cooling Unit 1) is marginally more primitive (with slightly higher Cr and lower Zr abundances) than the rest of the section below (cooling Units 2 and 3), although there is no systematic variation with depth (Fig. 3). Hole 802A, on the other hand, shows a chemical break in Zr content at 536 mbsf (Fig. 3) that corresponds to the boundary between cooling Units 14 and 15. Both the lower and upper chemical groups show a poorly defined trend towards higher Zr values with height in the "basement" (the highest value of 101 ppm Zr being at the top of cooling Unit 1). The significance of the Zr hiatus is not clear as other stable trace elements (incompatible and compatible) do not show this feature as strongly as Zr (Fig. 3) and the lavas above and below show broadly the same petrographic variation throughout.

The development of secondary minerals is a common feature of both Holes, and in particular, influences the distribution of most large-ionlithophile (LIL) elements (for example, Humphris and Thompson, 1978). Hole 800A alkalic dolerites are more hydrated (higher H_2O+) and oxidized (higher Fe_2O_3 /FeO ratio) than the mildly altered Hole 802A tholeiitic basalts (Fig. 4). Part of the high water content of the former group probably reflects the presence of extensive contact metamorphic biotite. Both groups exhibit considerable non-magmatic variation in their LIL-element contents and with ratios indicative of low-grade submarine alteration rather than the effect of magmatic processes; for example K/Rb ratios range from 400–1300 (Hole 800A dolerites) and 160–950 (Hole 802A basalts).

The Cretaceous alkalic dolerites (Hole 800A) and tholeiitic basalts (Hole 802A) exhibit the following primary chemical features (Figs. 5 and 6):

1. As might be expected all the alkalic dolerites are markedly enriched in incompatible elements (except Y) and exhibit different incompatible element ratios to the tholeiitic basalts. For example, alkalic dolerites have Zr contents >300 ppm and Zr/Y ratios of 13–16, whereas the tholeiitic basalts have about 60–70 ppm Zr and very low Zr/Y ratios between 2.5–2.8.

2. Both dolerites and basalts are characterized by very narrow ranges of incompatible elements such that both suites show little chemical variation (Fig. 5) and feature relatively uniform incompatible element ratios. Compatible elements, such as Cr and Ni, show a slightly greater diversity (Fig. 5) and together with a moderate range of FeO*/MgO ratios for dolerites (0.8–2.0) and basalts (1.0–1.6), indicate that only very minor mafic fractionation (olivine and clinopyroxene) has taken place throughout both groups. Compare, for example, the range in elemental compositions exhibited by tholeiitic lavas from the Nauru Basin and Ontong-Java plateau with that of Hole 802A tholeiites (Fig. 5).

3. Although chondrite-normalized rare earth element (REE) patterns (Fig. 6) are dissimilar for the alkalic dolerites and tholeiitic basalts, the range of variation and degree of REE fractionation is again very limited in each group. The dolerites show strong light REE enriched patterns, with $(La/Yb)_N$ from 17.5–20.0, whereas most of the basalts are characterized by relatively flat REE patterns ([La/Yb]_N 0.9–1.2) with about 10–12 times the chondritic abundances. Some of the slightly more fractionated samples (with lowest Cr and Ni) exhibit minor light REE enrichment ([La/Yb]_N 1.5–2.0) that is a function of clinopyroxene fractionation. However, the general similarity and parallelism of the REE patterns in both dolerites and basalts reflects the precipitation of phases that do not fractionate the REE to any degree and conforms with the observed presence of olivine and plagioclase as phenocrystic phases.

In general terms, the incompatible element-enriched features of the alkalic dolerites are characteristic of ocean island basalts (OIB). Although the low abundances of these elements and their distinctive ratios in the tholeiitic basalts are akin to some types of mid-ocean ridge basalt (MORB) (Bougault et al., 1980; Saunders, 1984), their flat REE patterns distinguishes them from typical N-type MORB. These distinctive chemical features find their counterpart in clearly different isotopic ratios, which implies the two groups were derived from different mantle sources (Castillo et al., this volume). This is substantiated by markedly different ratios of highly incompatible elements (La, Nb, Ta, Th) that reflect the nature of the source rather than variable mantle melting (Bougault et al., 1979). For example, La/Th and La/Ta ratios are consistently higher in the basalts (about 15 and 18, respectively) than the dolerites (about 10 and 12.5, respectively).

ERUPTIVE ENVIRONMENT AND ORIGIN

In this section we consider the general geochemical features of the dolerites and basalts in an attempt to determine their tectonic setting relative to other basic submarine volcanics, especially those considered part of the widespread mid-Cretaceous volcanic event. In terms of age and postulated eruptive setting, both Hole 800A dolerites and Hole 802A basalts are strong contenders for additional expressions of mid-Cretaceous volcanism (Lancelot, Larson, et al., 1990).

Alkalic Dolerite Sills

Thick massive doleritic units, generally interpreted as sills, have been drilled in a number of deep sea locations in both the Atlantic and Pacific oceans (for example, Leg 43, Tucholke, Vogt, et al., 1979; Leg 58, deVries Klein, Kobayashi, et al., 1980; Leg 61, Larson, Schlanger, et al., 1981; Leg 64, Curray, Moore, et al., 1982), although with a few exceptions most have tholeiitic compositions. Sills forming part of the oceanic crust of small basins, such as the Gulf of California (Saunders et al., 1982) and the Daito and Shikoku basins in the western Pacific (Marsh et al., 1980) are mainly tholeiites, although in a few cases relatively abundant alkalic basalt units may be present, such as in Hole 446A (Daito Basin). In the deeper water, open ocean environment, such as the Nauru Basin, the upper portion of the basaltic sequence constitutes a series of sills, although again they are tholeiitic in composition (Floyd, 1989). In some cases submarine sills appear to be closely associated with seamounts or ocean islands. For example, alkalic sills with high incompatible element contents (200-400 Zr ppm) were recovered from Leg 43 (Houghton, 1979), some of which were developed on the flanks of the New England Seamounts. In the Central Pacific Basin various doleritic rocks (some considered to be sills) with an alkalic parentage and containing primary brown amphibole and biotite, were probably derived from the nearby Line Islands or adjacent seamounts (Bass et al., 1973). It is apparent that the alkalic composition of some sills and their relatively close geographic position to nearby seamounts and ocean islands suggests they may be structurally related and form minor intrusions on their flanks. A seamount model based on the dissected La Palma Seamount Series (Canary Islands) suggests that small sill-like intrusions are a feature of early development and may well extend out beyond the seamount base (Staudigel and Schmincke, 1984).

As there is little chemical data on submarine Cretaceous-aged alkalic sills, we have compared the Hole 800A sills with oceanic islands, seamounts, and submarine sills of different ages but of alkalic origin. As seen in Figure 7, alkalic basalts from submarine sills show good covariance for individual suites, with Ti/Zr ratios ranging between 120 and 59. Note that the upper sequence Jurassic alkalic dolerites from Site 801 also plot within this range. Probably because of the opportunity to sample a wider range of material, oceanic island alkalic basalts, on the other hand, show a far wider range of TiO2 and Zr abundances, with a TiO2 peak that roughly separates the strictly basaltic compositions from the more evolved basic differentiates. The Hole 800A dolerites also show a minor peaked pattern (due to titanomagnetite fractionation), but differ from the "main alkalic array" with very low Ti/Zr ratios (about 36). In comparison the apparently evolved compositions of Hole 800A dolerites are also supported by the petrography (see above) which indicates that they are mainly of hawaiite composition. The dolerites also differ from other OIB in exhibiting highly fractionated REE patterns (Fig. 6) with (La/Yb)_N ratios of between 18-21. As shown in Figure 8, these sills are more

fractionated than most OIB alkalic rocks and are only matched by some of the Society Seamount lavas (Devey et al., 1990). Other incompatible element ratios are also distinctive, such as high Zr/Y (about 14) and La/Th (about 10) relative to other OIB, as well as to the Site 801 alkalic dolerites.

On the basis of the above data, the Site 800 alkalic dolerites have a chemical identity of their own, but are more akin to evolved alkalic OIB from islands and seamounts than alkalic basalts that appear to form part of the oceanic basement (for example Site 801). The particular chemistry of the alkalic dolerite (especially the high (La/Yb)_N and Zr/Y ratios) may suggest that their parental composition was generated by smaller degrees of partial melting relative to some of the other OIB, or that the source was generally more enriched in some incompatible elements. Site 800 is surrounded by a number of seamounts including Himu and Golden Dragon, both of which have isotopically distinctive compositions with a high proportion of the HIMU mantle end member and are almost isotopically identical to the Site 800 alkalic dolerites (Castillo et al, this volume). A HIMU component is also suggested by the specific chemical composition of the dolerites, such as the low Ba/Nb ratio (c. 5-8). As demonstrated by Sun and McDonough (1989) low Ba/Nb ratios correlate with low 87Sr/86Sr ratios that are generally typical of OIB with a HIMU component relative to other OIB with enriched mantle (EMI and EMII) components.

Both chemical and isotopic parameters suggest the dolerite sills are tapping the same mantle source as some of the nearby seamounts. In view of the seamount-sill relationships suggested elsewhere, Site 800 sills may also be volcanologically and genetically related, representing deep intrusive offshoots from the seamount flanks. Site 800 is surrounded by seamounts between 50–120 km away (Fig. 2, Lancelot, Larson, et al., 1990) and well within the distance of possible volcanic activity associated with active ocean islands. For example, the widespread alkalic volcanism associated with the Hawaiian South Arch (Lipman et al., 1989) occurs c. 200 km to the south of Hawaii Island, and in a similar manner the Site 800 sills may represent leakage of flank melts from plumes feeding the adjacent seamounts.

Tholeiitic Basalt Lavas

The extrusive pillow lavas of Site 802 with an age of about 115 Ma are probably a deep-sea expression of the mid-Cretaceous volcanic event, especially as they are about 50 m.y. younger than that predicted for the basement (mid-Jurassic) in this part of the Mariana Basin (Lancelot, Larson, et al., 1990). Apart from numerous seamounts generated in mainly Aptian times, large-scale volcanic provinces in the western Pacific generated during the mid-Cretaceous include Nauru Basin and various oceanic plateaus, such as Ontong-Java, Manihiki, Hess Rise etc. (Winterer, 1973; Schlanger et al., 1981). These submarine provinces are generally considered to represent episodes of crustal thickening soon after or as a continuation of axial activity (Saunders, 1986; Mahoney, 1987) generated above the inflated heads of large mantle plumes during the early Cretaceous (Richards et al., 1989). The tholeiitic basalts of the oceanic plateaus have a broad MORB-like geochemistry, although they exhibit a number of distinctive features that set them apart (Batiza et al., 1980; Floyd, 1989). In particular they are enriched in LIL elements, are characterized by generally flat (or less depleted) REE patterns and have different incompatible element ratios, in particular lower Zr/Nb, Zr/Ta, La/Ta, and Hf/Th, relative to normal-type mid-ocean ridge basalt (N-MORB).

Chemical tectonic environment discrimination diagrams, such as Hf-Ta-Th (Wood, 1980), Nb-Y-Zr (Meschede, 1986), and Zr/Y-Zr (Pearce and Norry, 1979), all group the Site 802 basalts as N-MORB. However, in this case, these diagrams are unreliable as the basalts are characterized by (1) flat REE patterns (Fig. 6) and (2) normalized multi-element distributions exhibiting LIL element and Nb-Ta enrichment (Fig. 9) relative to N-MORB. Highly incompatible element ratios are also different (Fig. 10) to Jurassic N-MORB (Site 801) with

Table 1. Major oxide and trace element data for Hole 800A alkalic dolerite sills.

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$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Depth (mbcf):	17-24	40-45	400.36	12-17	47-31	502.06	500 05	500.20	500.35	500 50	500.05	510.39	516.64
Mage Wate Wate <th< th=""><th>Cooling unit:</th><th>498.17 A1</th><th>498.40 Al</th><th>499.30 Al</th><th>A1</th><th>A2</th><th>A3</th><th>A3</th><th>A3</th><th>A3</th><th>A3</th><th>A3</th><th>A3</th><th>A3</th></th<>	Cooling unit:	498.17 A1	498.40 Al	499.30 Al	A1	A2	A3							
SiO ₂ 49.56 45.72 49.09 49.34 99.88 50.60 48.94 50.40 50.52 50.23 51.21 51.22 49.73 AL, 51.02 117 1.05 2.01 2.13 2.13 2.13 2.23 2.23 2.03 2.04 2.14 2.14 2.17 AL, 50, 15.86 14.77 15.97 15.91 15.61 15.65 15.89 15.83 15.60 16.15 15.96 16.23 16.56 15.24 16.56 15.25 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.56 15.26 16.26 16.56 15.26 16.26 16.56 15.26 16.26 16.56 15.26 16.26 16.56 15.26 16.26 16.56 15.26 16.26 16.56 16.26 16.56 16.26 16.26 16.56 16.26	Major oxides (wi	t%)												
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	SiO ₂	49.56	45.72	49.09	49.34	49.88	50.60	48.94	50.40	50.52	50.32	51.21	51.22	49.73
Al-O, 15.86 14.77 15.97 15.91 15.69 15.89 15.83 16.00 16.15 15.96 16.22 16.56 15.22 16.56 16.23 16.56 15.22 16.56 15.22 16.26 16.56 15.22 16.26 16.26 16.20 16.15 15.25 16.26 16.21 16.26 16.20 16.15 15.25 16.25	TiO ₂	2.17	1.95	2.01	2.19	2.21	2.13	2.11	2.13	2.23	2.03	2.04	2.14	2.17
Fe,O, 9.01 6.84 5.63 4.51 5.55 4.58 5.72 4.89 4.71 6.67 4.87 4.14 4.50 MaO 0.15 0.12 0.07 0.15 0.15 0.19 0.20 0.15 0.18 0.18 0.18 0.18 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.21 0.16 0.19 0.20 0.15 0.18 0.18 0.17 0.15 0.19 0.22 0.26 0.25 0.25 0.25 0.25 0.25 0.25 0.25 0.25	Al ₂ O ₃	15.86	14.77	15.97	15.91	15.69	15.89	15.83	16.00	16.15	15.96	16.22	16.56	15.92
FeO 0.37 1.06 2.82 4.92 4.01 5.12 3.78 4.86 4.60 3.38 4.30 4.84 5.30 MeO 1.5 0.12 0.07 0.15 0.12 0.07 0.15 0.15 0.19 0.20 0.15 0.18 0.16 0.19 0.21 0.14 0.19 0.25 0.55 0.54 0.62 0.54 0.65 0.55 0.55 0.55 0.55 0.55 0.55 0.55	Fe ₂ O ₃	9.01	6.84	5.63	4.51	5.55	4.58	5.72	4.89	4.71	6.47	4.87	4.14	4.50
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	FeO	0.37	1.06	2.82	4.92	4.01	5.12	3.78	4.86	4.60	3.38	4.30	4.84	5.30
	MnO	0.15	0.12	0.07	0.15	0.15	0.19	0.20	0.15	0.18	0.16	0.19	0.21	0.16
Caol 3.96 10.27 4.83 6.00 4.53 5.92 6.34 4.89 4.74 3.92 4.75 5.20 5.44 Na ₀ 2.266 3.63 2.94 3.92 3.86 4.12 4.07 3.87 4.05 3.79 4.07 4.41 4.22 K ₁ O 5.64 3.69 3.22 2.25 0.62 0.54 0.55 0.55 0.56 0.64 0.63 0.48 0.55 0.55 0.55 0.56 0.54 0.62 0.54 0.55 0.55 0.56 0.46 0.63 0.48 0.55 0.55 0.56 0.44 0.53 0.53 0.56 0.46 0.63 0.48 0.55 0.57 0.50 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.62 0.54 0.63 0.48 0.55 0.55 0.56 0.64 0.83 0.54 0.64 0.83 0.56 0.46 0.83 0.54 0.64 0.83 0.54 0.64 0.83 0.54 0.64 0.84 0.55 0.55 0.56 0.64 0.83 0.34 0.16 0.10 0.04 0.99.89 0.05 0.54 0.64 0.83 0.34 0.16 0.10 0.10 0.42 0.99.46 0.99.99 0.97.4 0.97.0 100.23 0.97.1 0.99.89 0.50 0.50 0.54 0.54 0.54 0.54 0.54 0.54	MgO	5.04	4.01	7.92	6.72	6.767	6.44	6.46	6.73	6.31	6.73	6.42	6.12	6.31
$ Na_{0} 0 266 3.63 2.944 3.92 3.86 4.12 4.07 3.87 4.05 3.79 4.07 4.41 4.22 Properties of the second s$	CaO	3.96	10.27	4.83	6.00	4.53	5.92	6.34	4.89	4.74	3.92	4.75	5.20	5.44
K ₂ O 5.64 3.69 3.22 2.55 0.62 0.54 0.55 0.53 0.56 2.03 2.22 2.24 1.94 (14) 0.55 0.55 0.55 0.56 0.54 0.62 0.54 0.55 0.53 0.56 2.04 0.64 0.63 0.48 0.55 0.56 0.54 0.64 0.55 0.53 0.56 0.44 0.55 0.56 0.44 0.55 0.56 0.54 0.64 0.55 0.53 0.56 0.44 0.55 0.56 0.44 0.42 0.42 0.43 0.46 0.42 0.42 0.43 0.46 0.42 0.42 0.43 0.46 0.42 0.42 0.43 0.44 0.42 0.43 0.45 0.44 0.42 0.43 0.45 0.44 0.44 0.44 0.44 0.44 0.44 0.44	Na ₂ O	2.66	3.63	2.94	3.92	3.86	4.12	4.07	3.87	4.05	3.79	4.07	4.41	4.22
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	K ₂ O	5.64	3.69	3.22	2.55	0.62	0.54	0.55	0.53	0.56	2.93	2.92	2.24	1.94
$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	P ₂ O ₅	0.55	0.55	0.56	0.54	0.62	0.54	0.55	0.53	0.56	0.46	0.63	0.48	0.55
	LOI	4.94	8.04	5.30	3.49	3.21	2.63	3.50	2.77	2.93	3.54	2.62	2.15	3.65
$ H_{0}O_{2} \begin{array}{ccccccccccccccccccccccccccccccccccc$	Total	99.92	100.64	100.38	100.26	99.36	100.42	99.46	99.99	99.74	99.70	100.23	99.71	99.89
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	H ₂ O+	3.38	2.21	3.61	2 38	3.04	213	2.50	2.41	2.49	3.14	2.22	1.97	2.70
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	CO ₂	1.34	5.10	1.38	1.18	0.22	0.60	0.64	0.38	0.34	0.18	0.16	0.12	0.43
Ba 475 251 339 337 388 317 368 359 402 421 413 316 391 Cc 95 74 109 81 104 80 95 91 80 111 81 88 90 Cr 203 140 127 134 80 155 81 120 82 81 93 326 38 Ga 61 22 23 22 22 24 24 24 23 24 21 31 34 21 14 44 43 61 29 42 39 Nb 52 48 50 51 53 53 51 55 65 40 48 38 42 43 41 46 44 39 43 42 39 35 15 55 54 49 45 65 40 48 38 34 24 31 23 22 13 14 14 39 43 42	Trace elements b	y XRF (pp	m)											
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ba	475	251	339	337	388	317	368	359	402	421	413	316	391
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ce	95	74	109	81	104	80	95	91	80	111	81	88	90
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Cr	203	140	127	134	80	155	81	120	82	81	93	326	38
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Cu	63	30	65	54	24	36	83	74	42	43	65	64	38
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Ga	21	22	23	22	22	22	24	24	22	24	23	24	21
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	La	48	31	54	31	52	31	41	44	43	61	29	42	39
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Nb	52	48	50	51	54	53	53	51	55	58	56	56	58
Ni 122 83 82 42 43 41 46 44 39 43 43 42 39 Pb 3 4 1 1 1 2 2 2 1 3 3 1 1 Rb 93 82 20 31 22 19 24 25 23 27 29 28 21 S 123 128 1054 273 263 332 421 315 305 268 274 271 350 Sr 340 466 493 566 525 608 578 548 544 473 526 597 574 V 165 188 187 183 186 192 187 185 190 154 211 189 Zr 23 23 23 23 360 390 362 344 379 390 400 394 386 Trace elements by INAA (ppm) Cs - 53 3.1<	Nd	56	37	60	27	58	35	45	49	45	65	40	48	38
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Ni	122	83	82	42	43	41	46	44	39	43	43	42	39
Rb 93 82 20 31 22 19 24 25 23 27 29 28 21 Sr 123 128 1054 273 263 332 421 315 305 268 274 271 250 Sr 340 466 493 566 525 608 578 548 544 473 526 597 574 V 165 188 185 187 183 186 192 187 185 190 154 211 189 Y 22 25 23 27 27 26 25 25 24 25	Pb	3	4	1	1	1	2	2	2	1	3	3	1	1
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Rb	93	82	20	31	22	19	24	25	23	27	29	28	21
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	S	123	128	1054	273	263	332	421	315	305	268	274	271	350
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Sr	340	466	493	566	525	608	578	548	544	473	526	597	574
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	V	165	188	185	187	183	186	192	187	185	190	154	211	189
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Y Ze	22	25	23	27	27	26	25	25	25	24	25	25	25
Zr 348 330 340 356 360 390 362 344 379 390 400 394 380 Trace elements by INAA (ppm) Cs — 3.53 0.14 0.15 0.15 — — 0.10 0.06 0.10 — 0.10 — Hf — 7.51 7.12 7.70 8.10 — — 8.10 8.29 8.19 — 8.30 — Sc — 16.3 17.8 20.3 18.6 — — 19.1 18.0 17.1 — 19.3 … Ta — 3.18 3.21 3.26 3.33 — — 3.19 3.53 3.71 — 3.63 — U — 1.93 1.58 1.25 1.23 — — 1.23 1.37 1.48 — 1.39 — Rare carth elements (ppm) La — 36.00 43.06 46.07 — — 43.60 41.40 39.00 — 45.99 — </td <td>Zn</td> <td>157</td> <td>80</td> <td>144</td> <td>11</td> <td>101</td> <td>12</td> <td>117</td> <td>82</td> <td>104</td> <td>114</td> <td>90</td> <td>90</td> <td>115</td>	Zn	157	80	144	11	101	12	117	82	104	114	90	90	115
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Zſ	348	330	340	330	300	390	302	344	379	390	400	394	380
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Trace elements b	y INAA (p	pm)	1000	21012						0.10		0.10	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Cs		3.53	0.14	0.15	0.15		-	0.10	0.06	0.10		0.10	-
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ht		7.51	7.12	7.70	8.10		_	8.10	8.29	8.19	_	8.30	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Sc	_	16.3	17.8	20.3	18.6			19.1	18.0	17.1		19.3	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	1a Th	_	3.18	3.21	3.20	3.33	_	_	3.19	3.53	3.71		3.03	
Rare earth elements (ppm) La - - 43.60 41.40 39.00 - 45.99 - Ce - 79.80 78.30 87.32 93.87 - - 43.60 41.40 39.00 - 45.99 - Ce - 79.80 78.30 87.32 93.87 - - 89.00 86.60 80.40 - 92.56 - Pr - - - 9.90 10.81 - - 10.06 - - 10.53 - Nd - 36.00 37.00 40.25 43.77 - - 40.34 38.00 37.00 41.51 - Sm - 7.96 7.75 7.96 8.4 - - 7.80 8.34 7.58 - 7.85 - Eu - 2.48 2.46 2.57 2.72 - 2.58 2.64 2.38 - 2.57 - - - - - 6.87 - - - -	In	_	1.93	3.04	4.20	4.30	_	_	4.10	4.00	1.48	_	1.39	_
Rare earth elements (ppm)La—37.1036.50 43.06 46.07 ——43.60 41.40 39.00 — 45.99 —Ce—79.8078.3087.32 93.87 ——89.00 86.60 80.40 — 92.56 —Pr———9.90 10.81 ——10.06———10.53—Nd—36.0037.00 40.25 43.77 —— 40.34 38.00 37.00 — 41.51 —Sm—7.967.757.96 8.4 ——7.80 8.34 7.58—7.85—Eu—2.482.462.572.72——2.582.642.82.65—Gd———6.847.44——6.87————Dy———6.847.44——6.87————Dy———6.847.44——6.87—————Dy———0.9460.87———5.27———5.37—Er———0.860.91——0.866———0.88—Er———0.860.91——0.866<			1.95	1.50	1.2.0	1.2.5			1.40		1.10			
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Rare earth eleme	nts (ppm)			15.04	14.00			10.00	41.40	20.00		10.00	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	La	-	37.10	36.50	43.06	46.07			43.60	41.40	39.00		45.99	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ce De		79.80	78.30	87.32	93.87			89.00	80.00	80.40		92.50	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Pr	-			9.90	10.81			10.06	20.00	22.00	_	10.55	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	ING See	-	36.00	37.00	40.25	43.77	_		40.54	38.00	37.00	_	41.51	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Sm		7.96	1.15	7.96	8.4			7.80	8.54	7.58		7.85	-
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Ed	-	2.48	2.40	2.57	2.12			2.38	2.04	2.38	_	6.95	577).
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Th	_	0.04	0.87	0.84	7.44	_	_	0.87	0.045	0.000		0.85	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Du		0.94	0.87	6.33	5.55	_	-	5.07	0.943	0.002		5 37	
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Ho	_			0.04	0.01	_	_	0.96				0.89	
In In <th< td=""><td>Fr</td><td>_</td><td></td><td>_</td><td>2.25</td><td>2.36</td><td></td><td>_</td><td>2.25</td><td></td><td></td><td></td><td>2 36</td><td>-</td></th<>	Fr	_		_	2.25	2.36		_	2.25				2 36	-
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Yh		1.44	1.20	1.67	1.50			1.55	1.50	1.48		1.67	
	Lu		0.197	0.177	0.26	0.24		_	0.24	0.204	0.201		0.26	

Note: Dash (----) indicates element not determined.

values intermediate to more enriched types, such as transitional MORB from the FAMOUS area of the Mid-Atlantic Ridge. Note that the only Pacific MORB with Zr/Ta ratios similar to the Site 802 basalts is from a failed rift (the Mendoza Rise), off the East Pacific Rise, drilled during Leg 92 (Leinen, Rea, et al., 1986). Although the basalts from Site 597 of this Leg have many depleted MORB-type characteristics (Pearce et al., 1986) they are very different to most Pacific and Atlantic N-MORB in terms of ratios of highly incompatible elements as seen in Figure 10. However, in many cases incompatible

element ratios overlap or are similar in the Site 802 lavas to the Nauru Basin and Ontong-Java tholeiites. Compatible element contents are also generally similar at the same level of magmatic evolution, although the range of Cr and Ni contents (up to high values) is not seen at Site 802. Isotopic ratios (Sr, Nd, and Pb) are also comparable to the Nauru Basin basalts (Castillo et al., this volume).

Overall, the Site 802 pillow lavas are chemically similar to other large-volume mid-Cretaceous volcanic provinces in the Pacific Ocean and we suggest that they represent a further example of this widespread

Table 1 (continued).

60R-1	60R-1	60R-1	61R-1	61R-1
15-20	44-46	69-74	10-16	17-24
526.05	526.30	526.59	535.40	535.47
A3	A3	A3	A3	A3
				-
49.90	50.67	50.80	50.68	50.21
2.12	2.14	2.09	2.12	2.29
16.06	16.11	16.09	15.99	15.99
4.33	5.27	5.66	4.89	4.99
4.96	4.87	3.85	4.86	4.89
0.14	0.16	0.14	0.12	0.12
6.36	6.48	6.75	6.53	6.67
5.94	5.00	4.64	5.26	5.05
4.04	4.00	4.02	4.04	4.06
2.29	2.50	2.49	2.29	2.33
0.48	0.55	0.46	0.47	0.60
3.17	2.51	3.23	2.59	3.04
99.78	100.26	100.22	99.83	100.24
2.22	1.88	2.50	2.20	2.83
0.95	0.20	0.16	0.29	0.20
200	290	410	221	271
808	380	419	02	3/1
92	104	/1	92	40
22	107	64	333	48
32	22	49	22	24
22	23	24	21	21
41	44	31	47	30
52	54	22	52	22
48	54	34	48	41
43	44	44	4.5	39
2	3	-	1	1
25	27	24	25	23
194	334	263	159	168
282	559	518	572	542
181	1/8	184	200	194
24	27	24	25	26
76	95	111	85	111
357	370	379	356	363
0.11	0.09	0.07	0.10	0.11
7.80	8.03	7.87	8.50	7.83
17.4	19.4	17.4	20.6	15.6
3.36	3.45	3.46	3.40	3.50
4.40	3.99	3.98	4.40	4.06
1.25	1.43	1.36	1.22	1.27
43.90	39.00	36.00	42.37	40.40
89.16	81.80	76.20	85.60	85.40
9.95	1000		9.78	
39.71	40.00	36.00	38.00	41.00
7.52	7.96	7.51	7.42	8.36
2.50	2.53	2.43	2.51	2.60
6.57			6.69	
	0.950	0.883		0.964
		-	5.26	-
5.22				
5.22 0.85		_	0.86	
5.22 0.85 2.29			0.86 2.31	
5.22 0.85 2.29 1.63	 1.52	 1.42	0.86 2.31 1.63	 1.57

volcanic event. Isotopically, Site 801 lavas fall outside typical Pacific MORB compositions (Castillo et al., this volume) and are similar to large oceanic plateaus (Castillo et al., 1991; Mahoney and Spencer, 1991) with OIB-like compositions. They differ, however, in volcanological development and environment to the oceanic plateaus (that appreciably thicken the oceanic crust near spreading axes), in that the Site 802 basalts probably represent a thin veneer over the local Jurassic crust and were developed in an off-axis environment a considerable time after the local basement was formed. An alternative (but less likely) possibility is that Site 802 tholeiites actually represent



Figure 3. Distribution of Zr, Th, Cr (Hole 802A) and Zr (Hole 800A) with depth below sea floor (mbsf). The Zr hiatus in Hole 802A occurs between cooling Units 14 and 15, although this is not marked by other trace elements, such as Th and Cr.

enriched ocean crust in the East Mariana Basin, developed and focussed by the local rifting of pre-existing Jurassic crust (for example, Castillo et al., 1991). Whatever the case, the nature of the mantle was represented by an enriched MORB-source composition, rather than a typically depleted composition.

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Core, section Interval (cm)	57R-2 109-115	57R-3 1-6	57R-3 37-42	58R-1 9-13	58R-2 96-100	58R-2 18-23	58R-2 91-96	58R-2 99-102	58R-2 117-122	58R-3 23-28	58R-3 90-95	58R-3 134-140
Depth (mbsf)	509.19	509.61	509.97	516.09	516.96	517.68	518.41	518.49	518.67	519.23	519.90	520.34
Cooling unit	A1	Al	Al	A2	A2	A2	A2	A3	A4	A5	A5	A5
Major oxides (wt	%)											
SiO ₂	48.66	49.50	49.96	50.56	49.55	50,76	49.18	49.40	49.36	50.46	49.34	50.66
TiO	1.42	1.17	1.18	1.13	1.09	1.11	1.19	1.22	1.15	1.22	1.10	1.14
ALO.	16.53	14.80	14.50	14.45	14.21	14 20	14.27	14.39	14.38	14.78	14.25	14.35
Fe-O-	5.91	4.10	3.56	3.75	3.07	3.31	3.46	2.54	3.75	3 69	4 40	3.43
FeO	4 75	6.43	7.26	7 33	8 58	8 29	8.02	8.83	817	6.36	7.56	7.28
MnO	0.16	0.20	0.21	0.19	0.21	0.18	0.20	0.21	0.21	0.16	0.20	0.17
MgO	7.42	7.87	7 79	7.68	8.05	7.50	7.69	7.66	7.68	7 34	7 71	8.03
CaO	8 53	11.70	11.69	11.72	11.76	11.69	11.73	11.78	11.87	11.96	11.73	11.93
Na.O	2.78	2.65	2.57	2.10	2.00	2.01	2.10	1.95	1.95	2 33	2.04	1 99
KO	0.84	0.10	0.09	0.08	0.07	0.10	0.24	0.18	0.14	0.08	0.25	0.03
P O	0.10	0.04	0.03	0.03	0.07	0.03	0.03	0.02	0.03	0.00	0.02	0.10
LOI	2.70	1.17	1.02	1.00	0.03	0.03	1.00	0.86	1.08	1.50	1.09	0.90
			1.02	1.00	0.70	0.21		0.00	1100			
Total	99.79	99.73	99.86	100.091	99.61	100.13	99.11	99.03	99.77	99.90	99.69	100.10
H ₂ O+	2.05	1.11	0.63	0.57	0.54	0.44	0.64	0.54	0.90	0.65	0.65	0.51
CO ₂	0.58	0.12	0.22	0.18	0.10	0.10	0.16	0.11	0.13	0.62	0.15	0.06
Trace elements by	v XRF (ppm))										
Ba	56	10	26	9	16	32	11	20	14	18	12	28
Ce	17	12	10	8	25	12	26	10	7	7	8	3
Cr	167	156	138	125	150	131	151	155	159	134	141	136
Cu	162	157	155	153	153	159	148	157	150	156	144	161
Ga	20	15	17	155	17	21	146	15	16	16	19	16
1.9	3	2	3	2	5	4	2	4	2	2	2	2
Nb	7	4	4	3	4	4	4	4	4	4	Ã	5
Nd	7	4	5	16	23	13	25	10	8	6	10	5
Ni	99	103	100	06	08	05	06	00	101	100	99	90
Dh	00	105	100	50	50	55	50	4	5	4	6	2
Pb	4	3	2	0	0	0	6	+	4	1	8	ĩ
S S	224	400	275	180	252	202	270	284	222	217	272	242
3	234	400	275	189	252	302	279	284	322	217	272	242
Sr	101	98	104	102	99	97	98	95	94	240	251	363
V	423	357	354	359	332	342	346	338	333	349	331	303
Y Z	28	25	24	25	24	25	26	25	25	25	24	24
Zn	119	92	90	91	95	91	94	97	94	93	92	94
Zr	101	64	67	66	65	64	65	65	65	65	63	63
Trace element IN	AA (ppm)		2.22	0.10		0.10			0.15	0.045	0.10	
Cs	0.09		0.20	0.10		0.10		_	0.15	0.045	0.19	
Hr	2.30	_	1.76	1.90	_	1.60			1.76	1.80	1.64	
Sc	48.8		49.4	46.5		47.8			49.0	48.6	47.7	
Та	0.46		0.19	0.186		0.165			0.17	0.16	0.165	
Th	0.62		0.22	0.28		0.26	-		0.25	0.26	0.18	
U	0.17		0.40	0.076	-	0,09		-	0.50	0.097	0.50	
Rare earth element	nts (ppm)										12000 at 1111	
La	9.07		3.17	3.93	-	3.83	0		3.08	4.08	3.00	
Ce	19.88	-	8.60	9.93		10.02	_	-	8.50	10.31	8.00	1.00
Pr	2.65	-		1.46	_	1.46				1.51		
Nd	12.73		15.00	7.94		7.76	-	-	6.20	8.21	6.00	
Sm	3.19		2.32	2.25	_	2.23		-	2.22	2.38	2.25	
Eu	1.22		0.88	0.89	_	0.87		-	0.87	0.91	0.814	
Gd	4.16			3.34		3.28	-	-	_	3.50		
Tb		-	0.57			_			0.56	_	0.54	
Dy	4.61	-	-	4.08	-	4.08	-			4.17		
Ho	0.87			0.82		0.83	-	-	-	0.83		
E.	2.52			2.47		2.44	_	_		2.48		
Er	4.34											
Yb	2.32	_	2.37	2.49	_	2.38	-	-	2.41	2.51	2.28	

Note: Dash (----) indicates element not determined.

Table 2 (continued).

58R-4	58R-4	59R-1	59R-1	59R-1	59R-1	59R-1	59R-2	59R-2	59R-2	59R-2	59R-3	59R-3	59R-3
24-29	65-70	7-12	70-75	78-82	130-136	138-142	8-13	15-19	100-105	123-127	17-22	40-44	127-132
520.74	521.15	525.47	526.10	526.18	526.70	526.78	526.98	527.05	527.90	528.13	528.49	528.72	529.59
A5	A5	A7	A8	A8	A9	A9	A9	A9	A10	A11	A12	A13	A13
50.64	49.32	49.46	49.59	49.76	49.76	49.27	49.95	49.58	49.52	49.38	50.48	49.98	50.34
1.17	1.12	1.21	1.22	1.14	1.10	1.10	1.24	1.11	1.21	1.10	1.16	1.23	1.20
14.10	14.55	14.72	14.23	14.00	14.26	14.43	14.20	14.31	14.42	14.37	14.50	14.27	14.24
3.01	3.87	3.60	2.41	2.73	3.55	3.00	2.07	3 30	2.60	3.65	3.18	2.03	2.65
9.35	8.22	7.63	0.57	2.13	0.55	7.99	2.97	0.00	2.00	9.47	7.92	0.17	2.05
0.55	0.23	7.05	0.37	0.03	0.52	7.00	0.29	0.03	0.34	0.47	7.03	9.17	0.33
0.17	0.21	0.19	0.20	0.20	0.21	0.21	0.19	0.21	0.20	0.21	0.18	0.21	0.21
7.50	1.57	7.98	7.94	7.92	7.00	7.64	7.68	7.71	8.07	/.6/	7.51	/.6/	1.12
11.65	11.96	11.97	12.13	12.02	11.85	11.58	11.85	12.02	12.12	11.85	11.84	11.85	11.78
2.06	2.04	2.08	1.96	1.91	1.92	2.07	2.00	1.92	2.01	1.95	2.13	2.00	1.97
0.08	0.13	0.08	0.06	0.08	0.14	0.18	0.20	0.14	0.05	0.13	0.11	0.10	0.08
0.02	0.03	0.03	0.10	0.03	0.02	0.02	0.09	0.02	0.02	0.02	0.02	0.02	0.02
0.95	1.05	0.98	1.11	1.04	1.03	1.09	1.05	0.90	0.99	0.91	0.87	0.84	1.08
99.76	100.06	99.93	99.56	99.53	100.02	99.45	99.80	100.06	99.56	99.72	99.91	99.37	99.64
0.46	0.94	0.64	0.43	0.53	0.92	0.73	0.58	0.54	0.50	0.71	0.51	0.50	0.86
0.10	0.08	0.13	0.45	0.19	0.92	0.75	0.13	0.13	0.37	0.07	0.24	0.18	0.10
0.10	0.08	0.15	0.20	0.18	0.07	0.24	0.15	0.15	0.27	0.07	0.24	0.18	0.10
12	22		10	24			22	17	0	16	17	22	20
13	22	7	19	34	22	14	23	17	9	16	17	32	29
10	10	8	4	11	4	5	6	6	7	5	8	8	10
136	180	139	131	144	166	148	138	155	133	165	134	150	145
154	156	163	164	155	152	152	157	151	157	151	151	150	156
18	17	15	15	14	18	17	16	17	17	18	17	18	16
2	3	1	3	4	4	3	3	3	1	1	2	2	2
3	4	3	5	3	4	3	5	4	4	5	4	4	3
4	5	9	8	7	12	10	8	12	15	4	7	10	20
93	110	103	96	93	105	100	100	100	98	105	98	100	102
7	4	4	1	4	4	5	4	6	6	5	6	4	6
1	i.	1		1	1	1	4	4	1	2	2	2	2
315	330	203	244	240	245	255	240	263	246	250	407	266	264
06	05	100	07	249	245	255	249	203	240	259	407	200	204
240	95	100	226	98	95	100	90	90	99	90	100	93	93
340	307	373	220	344	345	350	549	338	340	349	340	340	340
25	25	25	25	25	25	25	25	25	25	25	26	25	25
90	97	97	93	88	96	100	94	95	90	97	92	95	93
63	65	66	64	65	64	65	64	65	64	64	64	63	63
	0.30	0.10			0.08				0.10	0.16			0.20
	1.83	1.50	-		1.65	\rightarrow			1.60	1.69			1.76
	49.0	47.6	-	_	48.1				47.6	48.9	\sim	_	48.5
_	0.17	0.17			0.177				0.168	0.19	_	_	0.17
-	0.23	0.24	_		0.22				0.26	0.20		-	0.20
	0.30	0.112			0.30	12-12			0.074	0.40	-		0.30
	0.50	0.112			0.50	_			0.074	0.40			00
	2.10	3.06			2.16				1 92	2.14			2 1 2
	9.16	10.29			5.10				4.80	9.14	1		9.15
-	8.30	10.28	_		8.10	-			10.82	8.70		-	8.30
		1.47							1.51				17.00
_	9.00	8.30			5.00				7.94	8.00	_		17.00
-		2.35			2.35	-			2.20	2.31	_		2.25
—		0.91			0.863				0.88	0.88			0.86
—	1.	3.34				-			3.15	1000	-	—	1
	0.57		-		0.567	5 <u></u> 7				0.56		—	0.60
_		4.32							4.05				
_	_	0.85							0.80		_	_	
	_	2.59							2.43			_	_
-	2.41	2.57			2.37		_		2.47	2.43			2.36
_	0.358	0.43			0.359		_		0.42	0.364			0.372
	1000000000												1000000-00-0

Table 2 (continued).

Core, section Interval (cm)	60R-1 11-16	60R-1 78-84	61R-1 32-37	61R-1 139-144	61R-2 14-21	62R-1 0-7	62R-1 112-118	62R-2 30-35	62R-2 45–50	62R-2 53–59	62R-3 12-18	62R-3 27-33
Depth (mbsf) Cooling unit	534.81 A14	535.48 A14	544.52 A15	545.59 A16	545.84 A16	550.60 A17	551.72 A17	552.36 A17	552.51 A17	552.59 A17	553.60 A17	553.75 A17
Major oxides (wt	%)											
SiO ₂	50.01	51.00	49.31	50.62	50.31	50.76	50.77	50.48	49.35	50.79	50.41	50.65
TiO,	1.11	1.20	1.19	1.09	1.22	1.17	1.06	1.12	1.28	1.09	1.08	1.08
Al ₂ O ₃	14.44	16.28	14.64	14.13	14.05	14.40	14.11	13.88	14.43	14.25	14.27	14.17
Fe ₂ O ₃	2.85	4.15	3.71	3.28	3.19	3.41	3.39	3.08	4.53	3.10	3.78	3.31
FeO	8.79	4.77	7.70	8.21	7.87	7.52	8.29	8.46	7.10	8.49	7.62	8.35
MnO	0.23	0.12	0.18	0.18	0.17	0.18	0.19	0.18	0.24	0.20	0.18	0.18
MgO	7.89	7.16	7.56	7.73	7.84	7.78	7,79	7.85	7.87	7.76	7.61	7.64
CaO	12.15	10.97	11.72	11.54	11.28	11.29	11.72	11.68	11.74	11.79	11.63	11.82
Na ₂ O	1.99	2.52	2.08	1.99	2.35	2.24	1.96	2.04	2.10	1.93	1.92	1.99
K ₂ Ô	0.04	0.08	0.25	0.09	0.10	0.09	0.08	0.08	0.17	0.07	0.07	0.06
P ₂ O ₅	0.02	0.04	0.03	0.02	0.03	0.03	0.02	0.09	0.03	0.02	0.02	0.02
LOI	0.89	1.27	1.10	0.95	1.14	1.07	0.84	1.08	1.41	0.83	1.18	0.94
Total	100.40	99.55	99.48	99.83	99.53	99.92	100.22	100.02	100.14	100.31	99.76	100.22
11.01	0.45	0.95	0.75	0.51	0.01	0.75	0.46	0.52	0.60	0.46	0.82	0.45
CO ₂	0.18	0.85	0.13	0.12	0.09	0.11	0.40	0.04	0.78	0.12	0.82	0.20
Trace elements b	v XRF (non	n)										
Ra	20	21	20	25	20	24	37	25	30	33	29	29
Ce	5	4	4	2	2	1	0	4	5	4	3	4
Cr	146	150	160	123	140	134	143	127	130	134	147	133
Cu	151	150	150	153	156	150	145	162	161	157	150	157
Ca	151	19	150	17	17	1.59	17	15	17	18	16	15
La	10	10	13	17	17	3	2	15	17	10	2	2
Nih	2	2	2	2	1	3	2	2	4	4	2	4
NU	3	4	15	5	5	4	4	4	4	16	5	
NG	4	100	15	15	5	1	10	5	06	10	00	9
DI	102	100	95	95	98	90	99	94	90	93	39	5
PD	8		0	3	2	5	0	2	3	5	3	5
RD	1	1	0	2	201	2	1	2	4	201	240	212
5	223	239	264	275	294	194	310	339	300	304	348	313
Sr	96	118	100	98	93	96	95	94	95	98	94	95
V	541	416	361	332	342	356	323	328	340	330	332	329
Y	24	23	26	24	25	25	25	23	23	24	23	25
Zn	90	101	101	95	90	91	85	88	90	88	94	92
Zr	63	70	66	64	64	64	62	62	63	63	62	63
Trace element IN	AA (ppm)					0.05	0.00			0.10	0.00	
Cs	0.17	0.20	0.16		0.30	0.05	0.30			0.10	0.20	
HI	1.08	1.95	1.70		1.74	1.60	1.03			1.60	1.75	
Sc	48.2	54.3	48.7		48.4	50.1	48.0		_	48.6	48.2	
la	0.18	0.19	0.18	_	0.171	0.18	0.17		_	0.186	0.18	-
Th U	0.27	0.26	0.28		0.25	0.24	0.19	_	_	0.27	0.23	_
U	0.00	0.00	0.097		0.40	0.070	0.10			01007	0100	
Rare earth eleme	nts (ppm)		and the second		G. (1994)	/ up takan	12100					
La	3.10	3.52	3.77		3.28	4.21	3.02		-	3.65	3.13	
Ce	8.00	9.20	9.53		8.30	10.17	8.10		-	9.40	7.90	0.00
Pr	1.11		1.34		and the second	1.54	and the	_	_	1.50	-	
Nd	7.00	19.00	7.76	-	7.00	7.94	9.00		\sim	7.76	8.00	
Sm	2.29	2.41	2.30		2.36	2.26	2.21		_	2.26	2.34	
Eu	0.87	0.90	0.92		0.90	0.89	0.85	-	\rightarrow	0.89	0.83	
Gd			3.33			3.37		_	\rightarrow	3.24	_	-
Tb	0.57	0.53	-		0.57		0.55	_	\rightarrow		0.54	
Dy		-	4.17	_		4.08			<u> </u>	4.11		—
Ho	-	_	0.82			0.83			\rightarrow	0.82	-	-
Er	<u> </u>	—	2.47	-	-	2.48		_		2.49	-	
Yb	2.32	1.96	2.44	-	2.48	2.46	2.30		—	2.49	2.34	
Lu	0.369	0.281	0.41	-	0.375	0.44	0.361	_	-	0.42	0.367	-



Figure 5. Distribution of Cr, Ni, TiO₂, Y, Nb, and Ce relative to Zr in Hole 800A dolerites and Hole 802A tholeiites. Apart from the clear differences in incompatible element contents (excepting Y) between the magma types, note the restricted chemical range illustrated, especially relative to tholeiites from the Nauru Basin (data compiled from Larson, Schlanger, et al., 1981; Moberly, Schlanger, et al., 1986) and Ontong-Java Plateau (Leg 130; M. Storey, pers. comm.).



Figure 6. Envelopes incorporating the range of chondrite-normalized REE patterns for Hole 800A alkalic dolerites and Hole 802A tholeiitic flows. Within each group fractionation of the REE is very limited; the tholeiites typically exhibit virtually flat patterns.



Figure 7. Distribution of TiO_2 and Zr in alkalic basalts from deep-sea submarine sills (envelopes) and oceanic islands (trends) compared with Hole 800A dolerite sills and Leg 129 Site 801 alkalic flows (Floyd and Castillo, this volume). Data sources: Holes 167 and 170 (Bass et al., 1973); Holes 382 and 385 (Houghton, 1979); Holes 446A (Saunders et al., 1982); Mathematician failed rift (Batiza and Vanko, 1985); St. Helena (Chaffey et al., 1989); Caroline Islands (Mattey, 1982) and Society Seamounts (Devey et al., 1990).



Figure 8. Degree of light REE fractionation (chondrite-normalized La/Yb ratio) and TiO_2 distributions for various oceanic alkalic basalts relative to Hole 800A alkalic dolerites. Note the marked difference between the Leg 129 Site 801 alkalic basalts and the Hole 800A samples. Data sources: Hawaii (Basaltic Volcanism Study Project, 1981); Mathematician failed rift (Batiza and Vanko, 1985); Azores (White et al., 1979); St. Helena (Chaffey et al., 1989); Society Seamounts (Devey et al., 1990).



Figure 9. N-MORB normalized multi-element patterns for typical Hole 8002A tholeiite (Sample 129-802A-62R-2, 53–59) and most fractionated variant (Sample 129-802A-57R-2, 109–115). Note the enrichment in LIL elements, together with Nb and Ta. N-MORB normalization factors from Sun and McDonough (1989).



Figure 10. La versus Th and Ta versus Zr plots showing the distribution of Hole 802A tholeiites relative to Jurassic N-MORB (Leg 129, Site 801; Floyd and Castillo, this volume), enriched MORB from the FAMOUS area of the Mid-Atlantic Ridge (Blanchard et al., 1976; Joron and Treuil, 1989), and the Nauru Basin (Saunders, 1984; Floyd, 1986). Mendoza failed rift data from Pearce et al. (1986).