12. GEOCHEMICAL CHARACTERISTICS OF REFRACTORY SILICATE MELT INCLUSIONS FROM LEG 140 DIABASES¹

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ABSTRACT

Geochemical data from plagioclase-hosted silicate melt inclusions from Leg 140, Hole 504B diabase dikes are reported. Hand-picked plagioclase grains were heated to 1260°–1280°C to remelt the glass inclusions and to infer trapping temperatures. The samples were then polished to expose the inclusions, which were analyzed by electron and ion microprobes. Inclusion compositions are mainly in equilibrium with the host plagioclase and are more depleted in incompatible elements than the host rock. Simple crystal-liquid equilibrium calculations show that the melt inclusions could have been in equilibrium with depleted abyssal peridotite diopsides, whereas whole-rock basalt compositions generally could not have been. The melt inclusions are significantly more depleted than normal (N-type) mid-ocean-ridge basalt (MORB) and are consistent with being produced by 8%–16% incremental or open-system melting with 2% residual porosity in the peridotite source. These magmas were formed during pressure-release melting of the mantle over a range of depths between 30 and 15 km.

INTRODUCTION

In recent years, the process of mantle melting at mid-ocean ridges to produce mid-ocean-ridge basalt (MORB) has received much attention from earth scientists. Constraints on the physical melting process have come from studies of basalts (Klein and Langmuir, 1987), abyssal peridotites (Dick et al., 1984; Johnson et al., 1990; Johnson and Dick, 1992), experimental petrologic data (McKenzie and Bickle, 1988), isotopes (McKenzie, 1985; Rubin and Macdougall, 1988), and laboratory physical properties studies (Riley and Kohlstedt, 1991; Waff and Faul, 1992). Synthesizing the results of these studies, one can conclude that melting takes place over a column of mantle extending over a depth range depending on compositional and thermal conditions, that small melt fractions can segregate from their peridotite source, and that this process has profound effects on the chemistry of the melts and residues produced. The culmination of this process, variously called fractional, dynamic, incremental, and open-system melting, is that small melt fractions move through the melting column and aggregate or homogenize en route to the surface, where they are erupted as MORB. Thus, even high-Mg primitive MORB are probably not primary melts, and the search for the primary melts has led to investigations of silicate melt inclusions (Sobolev and Dmitriev, 1989; Johnson, 1990; Sobolev and Shimizu, 1993).

Interest in the chemical composition of melt inclusions stems directly from their potential role as windows into the mantle. Refractory melt inclusions with compositions that could have been in equilibrium with mantle olivine of Fo_{89-91} were described by a number of workers (e.g., Donaldson and Brown, 1977; Dungan and Rhodes, 1978; Sobolev and Naumov, 1985; Falloon and Green, 1986). If melting occurs over a depth range, and if minerals crystallize at any given depth(s), then they are theoretically capable of trapping melt anywhere along the ascent path. These primary melt inclusions (those fractions of melt trapped in a growing crystal as a result of some process or mechanism(s) that interrupts the growth of a perfect crystal, such as surface cracks, defects, or non-uniform growth [Roedder, 1984]) could therefore record the composition(s) of any liquids in

which the host mineral experienced a growth phase. Thus, the working hypothesis for this study is that if sub-ridge melting is nearly fractional, then small, primary melt increments may be preserved as melt inclusions in minerals growing from the early-formed melt, leading to the prediction that melt inclusions may span a range of compositions, the average of which represents MORB.

In evaluating the composition of melt inclusions, the effect of interaction between trapped melt and host phase must be addressed. While the absolute concentrations of elements may have been affected by post-entrapment crystallization and reequilibration, the ratios of incompatible elements are unlikely to have changed significantly and therefore can be used to infer the composition of the source region. Interpretation of inclusions as primary melts requires that they be in equilibrium with mantle olivines. Satisfying this requirement alone, however, is not proof that an inclusion represents an unmodified primary melt. Post-entrapment crystallization of plagioclase could have significantly altered the melt composition without changing its Mg/ (Mg + Fe) and thus it would appear to be in equilibrium with mantle olivine. A possible manifestation of post-entrapment crystallization would be enrichment in incompatible elements such as Ti, Zr, and the rare earth elements (REE).

In this paper, we report major and trace element compositions of silicate melt inclusions contained in plagioclase crystals from diabases recovered in Hole 504B on Ocean Drilling Program (ODP) Legs 137 and 140 (Fig. 1). From these data, we discuss their potential role as indicators of source and primary melt compositions and, in some cases, of the melting process by which they were produced.

ANALYTICAL TECHNIQUES

Mineral Selection

Five samples were chosen from Hole 504B based on the abundance of the plagioclase phenocrysts, the lack of plagioclase alteration, and representativeness of the recovered diabase types. Plagioclase was separated from the crushed rock, and about 20 grains were selected from the coarsest size fraction (0.7–1.0 mm) of each separation to ensure that they were phenocrysts and not groundmass plagioclase.

High Temperature Microthermometry

Five plagioclase grains were heated to between 1260° and 1280°C for 4 hr and then were quenched in water. This technique has been used previously to homogenize inclusions in phenocrysts (Vicenzi,

¹ Erzinger, J., Becker, K., Dick, H.J.B., and Stokking, L.B. (Eds.), 1995. Proc. ODP, Sci. Results, 137/140: College Station, TX (Ocean Drilling Program).

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Figure 1. Location of Site 504.

1985; Hansteen, 1991; Coombs et al., 1991; Sinton et al., 1993). The temperature was controlled to within 2°C of the indicated temperature using a Deltec vertical quench furnace, Eurotherm temperature controllers, and Pt-Pt10%Rh thermocouples. A second thermocouple, placed less than 10 mm from plagioclase grains, was used to monitor their temperature. The furnace atmosphere was maintained at an oxygen fugacity similar to that experienced by the host basalts with a mixture of H_2 and CO_2 . Table 1 lists the samples and run temperatures; abbreviations will be used throughout the text.

After the plagioclase cooled, it was mounted in epoxy, polished, and examined with the back-scattered electron detector on the microprobe (Fig. 2). In all 1270° and 1280°C runs, the partially crystalline inclusions in the plagioclase were completely melted to form a homogenous liquid plus a gas bubble in some cases. At 1260°C most inclusions were completely melted, but some still contained olivine (<1 μ m). This suggests that the melts were originally trapped at ≥1260°C, but less than 1270°C. The melt inclusions and surrounding plagioclase were analyzed by ion microprobe at Woods Hole Oceanographic Institution and by electron microprobe at Oregon State University.

Electron Microprobe Analyses

All analyses were done on a Cameca SX-50 four-spectrometer microprobe at Oregon State University Electron Microprobe Laboratory. Measured concentrations were corrected for atomic number, absorption, and fluorescence effects by the ZAF method. Standards consisted of natural minerals and basalt glass from the Smithsonian reference collection (Jarosewich et al., 1980) and a commercially prepared standard block (Astimex Scientific LTD, 16 Edgewood Crescent, Toronto, Canada).

Plagioclase phenocrysts were analyzed using the following standards: Kakanui anorthoclase (USNM 133868) for Si and Na; Kakanui

Table 1. Samples and run temperatures.

Abbreviation Run		Temperature (°C)	Sample	Piece	Lith. unit
HRN-3	228	1280°	137-504B- 181M-2, 109-114	9A	210
	233	1270°			
	235	1260°			
			140-504B-		
HRN-7	236	1260°	191R-1, 34-40	7A	218
HRN-10	234	1260°	200R-1, 16-24	5A, 5B	227
HRN-14	237	1260°	205R-1, 95-102	23, 24	235
		1260°	206R-1, 9-38	3, 4A, 7, 9	235
HRN-22	238	1260°	226R-1, 134-140	24	260

Note: All samples from Hole 504B.



Figure 2. Back-scattered electron photograph of a representative silicate melt inclusion analyzed in this study. Epoxy, which surrounds the two plagioclase grains, and vesicles appear black. Three light gray inclusions are numbered corresponding to Tables 2, 3, and 4. Ion probe analysis pits within the inclusion are indicated by the arrows.

augite (USNM 122142) for Fe and Mg; Lake County, OR, labradorite (USNM 115900) for Ca and Al (Jarosewich et al., 1980); and Hohenfels Sanidine for K. Analytical conditions were 15 kV accelerating voltage, 30 nA probe current, and 10 s counting time with a focused beam. Glasses were analyzed using the following standards: Kakanui anorthoclase for Na; Kakanui augite for Mg; Durango fluorapatite for P; Hohenfels sanidine for K; Makaopuhi basalt glass for Ca, Fe, Al, Ti, and Si; an unknown pyroxmangite for Mn; and Tiebaghi chromite for Cr. Analytical conditions were 15-kV accelerating voltage, 50-nA beam current, defocused beam (approximately 30 µm), and 10 s counting times for all elements except Cr, which was counted for 20 s. Na was always analyzed first and showed no significant loss during analysis. Precision of this technique (based on multiple analyses of the Makaopuhi basalt glass) is reported elsewhere (Forsythe and Fisk, 1993). Glass in the experiments was analyzed seven to 12 times at several different locations within the sample, and the averages of these analyses are given in Table 2.

Ion Microprobe Analyses

All ion probe analyses were done on the Cameca IMS 3-f located at Woods Hole Oceanographic Institution. For the isotopes ³⁰Si, ⁴⁷Ti, ⁵¹V, ⁵²Cr, ⁸⁸Sr, ⁸⁹Y, ⁹⁰Zr, and ⁹³Nb, a primary beam of negatively charged oxygen ions with a current of approximately 0.2 nA and a net

Table 2. Major element analyses of silicate melt inclusions in plagioclase from Legs 137 and 140 diabase, Hole 504B.

Run	Sample	T°C	Incl.#	Unit	SiO ₂	TiO ₂	Al ₂ O ₃	FeO	MnO	MgO	CaO	Na ₂ O	K ₂ O	P ₂ O ₅	Cr ₂ O ₃	Total	Mg#
OP228	HRN-3	1280	1	210	50.80	0.38	18.80	7.08	0.14	8.00	12.70	1.90	0.09	0.03	0.06	99.98	0.668
OP228 [*]	HRN-3	1280	2	210	50.50	0.4	18.40	7.23	0.12	7.79	13.40	1.91	0.02	0.04	0.10	99.91	0.658
OP228*	HRN-3	1280	3	210	51.20	0.41	18.10	7.42	0.18	7.64	13.00	1.77	0.02	0.05	0	99.79	0.647
OP228	HRN-3	1280	3	210	50.50	0.48	18.10	7.32	0.13	7.71	13.00	1.74	0.03	0.05	0.03	99.09	0.652
OP228_	HRN-3	1280	3	210	51.10	0.45	18.00	7.42	0.15	7.73	13.00	1.70	0	0.04	0.09	99.68	0.650
OP228*	HRN-3	1280	4	210	51.50	0.44	18.30	7.36	0.09	7.52	13.00	2.02	0.01	0.03	0.02	100.29	0.645
OP233_	HRN-3	1270	6	210	50.39	0.33	18.77	8.66	0.16	5.74	11.99	2.03	0.01	0	0.02	98.10	0.542
OP233*	HRN-3	1270	7	210	49.83	0.37	18.37	9.01	0.14	5.80	12.32	2.58	0.02	0	0	98.44	0.534
OP233	HRN-3	1270	8	210	48.50	0.15	21.76	7.57	0.11	9.22	8.66	2.38	0.05	0.05	0	98.46	0.685
OP233	HRN-3	1270	9	210	51.60	0.36	17.88	4.65	0.18	8.83	14.46	1.87	0.02	0	0.02	99.87	0.772
OP233_	HRN-3	1270	10	210	50.74	0.47	17.63	3.62	0.16	9.14	14.56	1.81	0.04	0.04	0.10	98.30	0.818
OP235"	HRN-3	1260	11	210	53.27	0.89	18.76	4.22	0.17	7.92	11.69	2.60	0.05	0.03	0.01	99.62	0.770
OP235_	HRN-3	1260	12	210	47.78	0.74	17.92	9.47	0.21	6.80	12.95	2.30	0.01	0	0.02	98.19	0.561
OP236*	HRN-7	1260	1	218	50.63	0.43	17.63	4.54	0.15	9.31	13.95	1.60	0.01	0.05	0.10	98.40	0.785
OP236	HRN-7	1260	2	218	51.19	0.37	17.92	5.20	0.15	8.83	12.39	1.98	0.06	0	0.06	98.16	0.752
OP236	HRN-7	1260	3	218	49.48	0.47	17.64	8.41	0.18	7.29	12.99	2.12	0	0.04	0.06	98.68	0.607
OP236	HRN-7	1260	4	218	51.31	0.55	16.76	5.11	0.12	6.58	16.12	1.86	0.06	0.08	0.10	98.64	0.696
OP236	HRN-7	1260	5	218	51.24	0.61	16.60	5.19	0.14	6.46	16.12	1.86	0.06	0.01	0.08	98.37	0.689
OP234_	HRN-10	1260	1	227	49.41	0.39	18.46	4.08	0.20	10.73	12.55	1.99	0.05	0	0.04	97.90	0.824
OP234 [*]	HRN-10	1260	13	227	50.80	0.62	20.39	2.68	0.13	8.96	13.28	2.27	0.06	0.01	0	99.19	0.856
OP237	HRN-14	1260	1	235	48.83	0.81	17.36	8.03	0.17	6.86	12.97	2.24	0.01	0.08	0.04	97.41	0.604
OP237	HRN-14	1260	2	235	49.14	0.49	17.99	8.17	0.11	6.79	12.79	2.16	0.04	0.01	0.04	97.73	0.597
OP237	HRN-14	1260	3	235	52.19	1.18	17.67	4.82	0.23	6.75	13.08	2.42	0.05	0.03	0.03	98.45	0.714
OP237	HRN-14	1260	4	235	52.34	1.30	17.91	4.41	0.20	6.79	12.99	2.47	0.06	0.07	0.03	98.57	0.733
OP237	HRN-14	1260	5	235	50.69	0.52	18.07	5.97	0.18	8.46	12.99	2.12	0.05	0.05	0.04	99.15	0.716
OP237	HRN-14	1260	6	235	50.14	0.43	18.98	6.26	0.15	9.62	11.90	2.23	0.05	0.06	0.06	99.88	0.733
OP237	HRN-14	1260	7	235	51.52	0.46	17.81	5.54	0.15	7.53	12.86	2.27	0.08	0.08	0.04	98.34	0.708
OP237	HRN-14	1260	8	235	50.08	0.50	17.83	6.98	0.19	7.31	13.28	2.23	0.03	0	0.04	98.47	0.651
OP237	HRN-14	1260	9	235	52.43	1.08	17.92	5.02	0.23	6.92	12.87	2.53	0.05	0.06	0.05	99.15	0.711
OP238	HRN-22	1260	9	260	49.49	0.22	17.47	7.66	0.16	7.76	13.09	2.34	0.03	0.06	0.04	98.32	0.644
OP238	HRN-22	1260	10	260	50.04	0.57	17.82	6.95	0.15	7.95	12.52	2.19	0.03	0.11	0.06	98.39	0.671
OP238	HRN-22	1260	11	260	50.20	0.62	17.95	6.88	0.21	8.18	12.63	2.05	0.04	0.02	0.06	98.84	0.679
OP238_	HRN-22	1260	12	260	49.06	0.24	17.44	7.84	0.17	7.85	12.95	2.39	0.03	0	0.06	98.03	0.641
OP238	HRN-22	1260	13	260	48.99	0.17	17.20	5.05	0.18	9.69	12.74	1.80	0.04	0.09	0.04	95.99	0.774
OP238	HRN-22	1260	14	260	47.54	0.15	17.36	5.51	0.11	10.80	12.73	1.62	0	0.03	0.03	95.88	0.777
OP238*	HRN-22	1260	15	260	48.50	0.14	17.24	6.03	0.15	8.70	12.38	1.88	0.06	0.05	0.04	95.17	0.720
OP238	HRN-22	1260	16	260	50.69	1.20	18.80	9.65	0.24	3.89	9.65	3.56	0.06	0.09	0.02	97.85	0.418
OP238	HRN-22	1260	18	260	48.24	0.95	18.00	8.53	0.10	6.42	12.03	2.53	0.02	0.08	0.06	96.96	0.573

Note: *Inclusions for which ion microprobe analyses are reported in Table 3.

energy of 12.55–12.65 keV was focused to a spot 8–12 μ m in diameter. For rare earth element isotopes (REE) ¹³⁹La, ¹⁴⁰Ce, ¹⁴⁶Nd, ¹⁴⁷Sm, ¹⁵³Eu, ¹⁶³Dy, ¹⁶⁷Er, and ¹⁷⁴Yb, a primary beam current of 1–2 nA was focused to a 15–25 μ m spot, and a high energy offset of –50 eV was used. Positively charged secondary ions were mass analyzed by a double focusing mass spectrometer with a high energy offset of –90 eV and an energy bandpass of ±10 eV for a net secondary accelerating voltage of 4400–4420 eV, and were detected and counted by a 17-stage Allen-type electron multiplier in pulse counting mode.

Individual spot analyses included five cycles of measurements, each cycle beginning with backgrounds and stepping progressively up mass to previously located peaks. Background and deadtime corrected secondary ion intensities were normalized to either ¹⁶³Dy, for REE, or ³⁰Si for the other trace elements. Intensity ratios were converted to concentrations using empirically derived intensity ratioconcentration correction factors (working curves) determined from well-analyzed glass standards of similar Si composition. All glasses used as standards were previously analyzed for REE and other trace elements by isotope dilution mass spectrometry (ID) and X-ray fluo-rescence (XRF) (Bender et al., 1984).¹⁵³Eu analyses are less accurate than other REE due to isobaric barium oxide and hydroxide molecular ion interferences (Johnson and Dick, 1992).

Uncertainties in trace element measurements by ion microprobe are a function of uncertainty in slopes of working curves and from variations in counts in individual point analyses (counting statistics). Uncertainty in the working curve slope is itself a function of ion probe counting statistics and the quality of the standard (i.e., its homogeneity and the accuracy of its composition as determined by another analytical technique). Slope uncertainty influences the accuracy of the calculated unknown concentration and is minimized by using high quality standards, replicate measurements, and consistent operating conditions. Using standards with nearly the same major element concentrations as unknowns minimizes matrix effects. Slopes for each element were calculated based on normalized silica concentration in the standard to compensate for silicon-induced bias in the intensity ratios. Slopes and uncertainties calculated for this study are 3 relative percent (r.p.) for Ti, 5 r.p.–10 r.p. for V, Cr, Sr, Y, Zr, Ce, Nd, Sm, Dy, and Yb, 10 r.p.–15 r.p. for Nb, La, and Er, and >20 r.p. for Eu. Trace element concentrations for the melt inclusions are reported in Table 3.

RESULTS

Evaluating Inclusion Compositions

Measured inclusion compositions could have been affected by overlap of the ion beam onto plagioclase during ion probe analyses or by melting of the host plagioclase during homogenization. Extreme care was taken to avoid beam overlap onto the plagioclase. An independent assessment of the analyses' accuracy was performed by plotting the Ti contents of the inclusions as determined by both ion and electron microprobes. Ideally, data on this plot should define a slope of one. The results of this comparison are shown in Figure 3. Only inclusion 7 in HRN-3 at 1270°C plots significantly off the 1:1 line. Assuming that the electron probe analysis is good, based on that machine's superior optical imaging capabilities, a point below the line could be caused by the ion beam overlapping onto the host mineral, causing a dilution of Ti. Another explanation could be that the inclusions are heterogeneous or contain impurities, meaning that the inclusion is not a homogeneous glass and that differences in beam placement between electron and ion probe analyses could account for compositional disagreement. However, overlap of the ion beam onto plagioclase should cause a significant increase in the measured Sr content, which is not observed in this sample (Fig. 4), and the cause for the discrepancy in Ti measurements is not understood.

A second test of beam overlap or plagioclase melting was applied to the samples based on the relative compatibility of Sr and Eu in

Table 3. Trace element concentrations in Leg 137 and 140 melt inclusions (in ppm).

Run	Sample	T°C	Incl.#	Unit	Ti	v	Cr	Sr	Y	Zr	Nb	La	Ce	Nd	Sm	Eu	Dy	Er	Yb	Ti/Zr	(La/Sm) _n
OP228	HRN-3	1280	2	210	2243	165	288	46.6	9.2	14	0.15	0.19	0.81	1.13	0.55	0.25	0.95	0.55	0.57	160	0.22
OP228	HRN-3	1280	3	210	2457	189	292	42.5	10.6	15	0.23	0.35	1.21	2.15	0.94	0.37	2.03	0.88	1.06	164	0.23
OP228	HRN-3	1280	4	210	2355	179	283	43.6	10.2	15	0.16					156					
OP228	HRN-3	1280	5	210	3356	159	150	47.3	14.2	23	0.34	0.49	2.09	3.19	1.23	0.30	2.33	1.77	1.45	144	0.25
OP233	HRN-3	1270	7	210	983	128	535	1.2	3.2	4.1	0.55	0.21	0.65	1.74	0.66	0.19	0.80	0.80	0.77	239	0.20
OP235	HRN-3	1260	11	210	5302	171	185	46.8	20.2	35	1.77	0.70	3.26	4.81	2.34	0.55	4.29	2.89	3.07	152	0.19
OP236	HRN-7	1260	1	218	2619	176	360	46.9	17.7	28	0.87	0.60	2.05	4.30	2.20	0.61	3.46	2.57	2.30	95	0.17
OP234	HRN-10	1260	1	227	1888	134	247	93.2	10.1	13	1.05	0.42	2.00	2.93	1.11	0.80	2.16	1.37	1.12	146	0.24
OP234	HRN-10	1260	13	227	4056	193	198	71.4	15.2	36	1.32	1.06	3.64	5.24	2.26	0.85	3.81	2.65	2.85	113	0.29
OP237	HRN-14	1260	5	235	3081	172	223	61.3	13.9	22	1.30	0.72	3.15	4.09	1.74	0.63	3.3	1.94	2.16	142	0.26
OP238	HRN-22	1260	13	260	1122	173	355	19.7	15.4	3.0	0.78	0.73	3.54	5.29	2.89	0.78	4.28	2.47	1.96	369	0.16
OP238	HRN-22	1260	14	260	1081	183	332	40.1	10.2	4.0	0.71	0.12	0.42	1.10	0.79	0.19	1.00	0.64	0.94	270	0.10
OP238	HRN-22	1260	15	260	880	105	370	7.9	3.3	3.3	0.39	0.54	1.97	3.11	1.76	0.80	2.23	1.55	1.66	269	0.19



Figure 3. Comparison of Ti analyses by electron and ion microprobes to demonstrate agreement of the techniques. The significance of Sample HRN-3, inclusion 7, is discussed in the text.

plagioclase relative to surrounding elements on a typical compatibility diagram (spider diagram—e.g., see Fig. 9). To represent the abundance of these two elements relative to their adjacent elements, we calculated $Sr^* = (Ce_n + Nd_n)/2$ and $Eu^* = (3Sm_n + Dy_n)/4$, and calculated the ratios Sr/Sr^* and Eu/Eu^* . Values greater than unity mean that there is a positive concentration anomaly for that element. If both values are greater than one in a sample, then there is a possibility that significant beam overlap onto plagioclase occurred. As can be seen in Figure 4, two samples plot in the top right quadrant, and their compositional characteristics may be influenced by plagioclase contamination.

Post-entrapment Crystallization of Host Plagioclase

To discuss primary inclusion compositions, the extent of postentrapment interaction must be evaluated. Ways in which the inclusion composition might be altered after trapping include crystallization of the trapped melt, diffusive re-equilibration of the host phase and the melt inclusion at different pressures and temperatures than those prevailing at the time of trapping, and possible decompression melting of the host phase upon ascent. We do not expect that the pressure difference between entrapment and the surface has a significant effect on the composition of the melt inclusions, but both crystallization of the inclusion and melting of the host plagioclase will affect the composition of the inclusion.



Figure 4. Sr/Sr* vs. Eu/Eu* in the melt inclusions. These values are calculated based on chondrite normalized values of these elements relative to adjacent elements in the spider diagram and are used to assess ion beam overlap from inclusion onto host plagioclase. Since both of these elements are compatible in plagioclase and their surrounding elements (Ce and Nd, and Sm and Dy, respectively) are incompatible, beam overlap with plagioclase will result in a positive value for both of these ratios. Analyses of two samples, HRN-3 inclusion 2 and HRN-10 inclusion 1 appear to be contaminated by plagioclase.

We approached the question of post-entrapment interaction between host plagioclase and trapped melt in two ways. First, we assumed that the host plagioclase was in chemical equilibrium with the melt at the time of entrapment. Compositions of the host plagioclase and melt inclusions lie along an empirically determined relationship between Ca/(Ca + Na) ratio of the glass and of the plagioclase (Fig. 5; data from Falloon and Green, 1986, and Tormey et al., 1987), indicating that the rehomogenized melt inclusions from the experiments were in equilibrium with the host plagioclase. Second, melting of the host plagioclase (Anss-Anon) from 1260°-1280°C does not appear to be significant because the melting of the endmember plagioclase will increase the Al2O3 content of the melt about 0.25 wt% per °C above the plagioclase-olivine cotectic (Nielsen, 1988; 1990). The range of Al₂O₃ contents of the glass inclusions suggests that they were not heated more than about 10°C above the cotectic except for two inclusions with more than 20 wt% Al2O3. The trace elements Sr and Eu (discussed below) also indicate that little plagioclase was incorporated into the glass inclusions.

If the inclusions contained crystals of plagioclase that were not completely melted in the microthermometry experiments, the inclusion composition would lie above the line in Figure 5. The fact that residual crystals are extremely rare and only three inclusion analyses (in the most refractory plagioclase) lie a significant distance above the line indicates that in most cases the inclusions were melted completely.



Figure 5. Ca/(Ca + Na) ratios of melt inclusions vs. those of host plagioclase close to the inclusion. The line is the empirical equilibrium relationship between Ca/(Ca + Na) of melts and plagioclase derived from experimental data (regression line from data in Falloon and Green [1986] and Tormey et al. [1987]). Most of the inclusions plot within analytical error (shown by error bars) of the line indicating equilibrium between host and melt.

Melt inclusions vary significantly within single plagioclase grains and from grain to grain in the same rock, making it difficult to compare experiments at different temperatures. For example, the two inclusions at 1260°C in Figure 6A are from the same plagioclase grain. Even so, remelting of inclusions in sample HRN-3 at 1260°, 1270°, and 1280°C produced relatively constant Al₂O₃ compositions with increasing temperature (Fig. 6B). We take this to mean that the most refractory melt inclusions were trapped at about 1280°C.

Major Elements

Major element compositions of the inclusions are plotted on the plagioclase-olivine-silica face of the CaO-MgO-Al₂O₃-SiO₂ (CMAS) tetrahedron using the projection scheme of Walker et al. (1979) (Fig. 7). Also plotted on the diagram are liquidus isobars drawn through 5, 10, and 15 kbar experimental data points from the literature (Takahashi and Kushiro, 1983; Falloon and Green, 1987; Falloon and Green, 1988; Falloon et al., 1988; Kinzler and Grove, 1992). All of the data plot between the 5 and 10 kbar isobars, and roughly define trends oblique to them. There does not appear to be a strong dependence on run temperature, as indicated by HRN-3 data, but there is significant scatter in the data sets.

The range of $Mg^{2+}/(Mg^{2+} + Fe^{2+})$, referred to as Mg#, of the melt inclusions is 0.53 to 0.78 (Table 2); Mg# greater than 0.70 indicates that the inclusion could have been in equilibrium with mantle olivine of >Fo₈₉. This assumes that all iron is ferrous and the olivine-melt FeO-MgO exchange coefficient is 0.306 (Roeder and Emslie, 1970). Table 4 summarizes the samples whose Mg/(Mg + Fe) could be in equilibrium with mantle olivines.

Trace Elements

Rare earth elements in the melt inclusions are plotted in Figure 8 along with fields for basalts from Hole 504B and hypothetical melts in equilibrium with abyssal peridotites (Johnson et al., 1990; Johnson and Dick, 1992). Most of the samples show negative anomalies in Zr and Ti (Fig. 9), which were shown to be characteristic features in diopsides in abyssal peridotites (Johnson et al., 1990) as well as in melt inclusions from Mid-Atlantic Ridge basalts (Johnson, 1990; Sobolev and Shimizu, 1993).

DISCUSSION

The melting process at mid-ocean ridges has been modeled in many ways, but most geochemists now concur that it is an incremental pro-



Figure 6. Change of (A) SiO_2 and (B) Al_2O_3 content of melt inclusions in Sample HRN-3 with experimental run temperature. The convergence of SiO_2 and the constant Al_2O_3 may mean that, while the inclusions are remelting at increasing temperatures and subsolidus devitrification products enter the melt, the host plagioclase is not melting.



Figure 7. Plagioclase-olivine-quartz projection in the CMAS system for MORB inclusions using the projection algorithm of Walker et al. (1979). Labeled pseudocotectics at various pressures are drawn through experimental data of Takahashi and Kushiro (1983), Falloon and Green (1987, 1988), Falloon et al. (1988), and Kinzler and Grove (1992). Inclusions (symbols) define a trend oblique to the pseudocotectics, consistent with derivation by melting over a depth range from 30 to 15 km (10 to 5 kb).

Table 4. Measured Mg# in melt inclusions and calculated equilibrium forsterite (Fo) contents in mantle olivine.

Sample	Inclusion	Mg#inclusion	Equilibrium Fo			
HRN-3	9	0.772	0.019			
	10	0.818	0.937			
	11	0.770	0.918			
HRN-7	1	0.785	0.924			
HRN-10	1	0.824	0.940			
	13	0.856	0.952			
HRN-14	3	0.714	0.893			
	4	0.733	0.901			
	5	0.716	0.894			
	6	0.733	0.901			
	7	0.708	0.890			
	9	0.711	0.891			
HRN-22	13	0.774	0.919			
	14	0.777	0.921			
	15	0.720	0.896			

Note: $Mg# = Mg^{2+}/(Mg^{2+} + Fe^{2+})$.



Figure 8. Chondrite normalized rare earth element concentrations in inclusions other than the two that showed both Sr and Eu anomalies in Figure 4. Lightly shaded field is the range of basalt compositions from Hole 504B showing that the majority of the inclusions are more depleted than the host basalts. The field for hypothetical liquids in equilibrium with abyssal peridotite diopsides (Johnson et al., 1990) is shown in heavy shading. Assuming that the peridotites represent the final residue of melting at the top of the mantle melting column, liquids in equilibrium with those rocks should be extremely depleted in incompatible elements, while MORB plot outside this field. Many of the inclusions plot in the lower portion of the field.

cess with the generation and nearly simultaneous removal of small melt fractions (Langmuir et al., 1977; Maaløe, 1982; McKenzie, 1985; Rubin and Macdougall, 1988; Johnson et al., 1990; Johnson and Dick, 1992; Sobolev and Shimizu, 1993). While the models differ in detail, they are similar in that melt is efficiently removed from the system soon after formation with only small amounts retained in the residue at any one time. The model we employ in this paper, opensystem melting (Johnson and Dick, 1992), allows for melt formation, then rapid and efficient segregation and flow through overlying peridotites. Thus, the source for melting is peridotite + melt-from-below at all times, and the model allows for a melting column open at the bottom to infusion of melts. Using the open-system model, the amount of transiently retained melt in mid-ocean-ridge settings, based on trace element concentrations in peridotites, ranges from 0.2% to 3.0% (Johnson and Dick, 1992). This range implies segregation of very small melt fractions, nearly fractional melting in many cases.

If sub-ridge melting involves such small melt fractions, then it follows that late-stage, highly depleted, refractory melts must exist



Figure 9. **A.** Anomalies in non-rare earth trace elements Zr and Ti expressed as the ratios of chondrite-normalized Zr and Ti to interpolated value of these elements between the closest analyzed rare earth elements on an incompatibility (spider) diagram. **B.** Change in Zr/Zr* with Ti used as a relative measure of crystal fractionation or melting. Zr* is calculated as $(Nd_n + Sm_n)/2$ and Ti* is calculated as $(3Sm_n + 5Dy_n)/8$. Thus, Zr/Zr* and Ti/Ti* are measures of the anomalies of these elements on spider diagrams, with values less than 1 indicating negative anomalies, which can be seen in most of the inclusions. Shaded field represents hypothetical melts in equilibrium with abyssal peridotite diopsides (Johnson et al., 1990; Johnson and Dick, 1992). If the peridotites represent the final residue of melting over a melting column, then the inclusion compositions are consistent with their being the complementary melt increments from this process. Arrows marked "m" and "c" are the trajectories expected for fractional melt increments and crystallization, respectively. Normalizing values from Anders and Grevesse (1989).

between the point of magma segregation and the region of aggregation and final mixing prior to eruption as MORB. Such melts are a necessary consequence of low melt retention open-system melting, and their recognition is of critical importance in understanding processes of MORB petrogenesis in general, and in evaluating the validity of the open-system melting model in particular. These melts will rarely reach the surface unaffected by mixing, making their existence difficult to directly document. However, minerals growing in a magma occasionally trap small drops of melt on their growing surface or in their skeletal interiors (Roedder, 1984) and preserve them as melt inclusions, more or less unaffected by external processes.



Figure 10. Inclusion compositions in terms of Ti/Zr and $(La/Sm)_n$. The open system melt model discussed in the text is plotted as the solid curve with pluses every one percent melting using the parameters in Table 5. The dashed curve is the second stage of a two-stage batch melting process after an initial 15% batch melting (star symbol); square symbols are at 5% melt intervals. Most of the inclusions plot outside the field of 504B basalts and are consistent with 8%–16% open-system melting with a constant 2% melt retained in the residue throughout the melting event.

Trace element concentrations in the melt inclusions are generally more depleted than associated basalts from Hole 504B (Figs. 8 and 9). Furthermore, the inclusions display large negative anomalies in Ti and Zr, which are uncommon in 504B basalts (Fig. 9A). Although the inclusions are more depleted than associated basalts, they lie within the range of values for melts in equilibrium with diopsides from abyssal peridotites (Figs. 8 and 9). Assuming that the peridotites represent the final residue of melting over a melting column, the inclusion compositions are consistent with their being the complementary melt increments from this process. Furthermore, the inclusions cannot be related to most of the analyzed 504B basalts by simple crystal fractionation ("c" in Fig. 9).

Melting Model

Ti/Zr ratios in MORB are generally between 80 and 125. This range is roughly centered about the chondritic ratio of ~110 and is consistent with the hypothesis that MORB represent aggregated melts produced by incremental melting of the mantle. Trace element concentrations in abyssal peridotites (Johnson et al., 1990) argue against the former process. During fractional melting, individual melt fractions will not be constrained to asymptotically approach the Ti/Zr ratio of their source, as will batch and integrated melts, but instead will reach much higher values, the limit being infinity for advanced increments of perfect fractional fusion. This means that Ti/Zr ratios in both incremental or fractional melts and their residues can increase to greater than the ratio in their source as fractional melting proceeds.

If the upper mantle source for MORB has a roughly chondritic Ti/Zr ratio, then melts with Ti/Zr >> chondritic could be produced from a source that has been depleted by varying degrees of open-system melting. To evaluate the inclusion compositions reported here in light of this model, we have plotted Ti/Zr vs. (La/Sm)_n in the melt inclusions in Figure 10. Also plotted are theoretical curves calculated using the open-system model presented by Johnson and Dick (1992) and Johnson and Kong (1992) using the parameters listed in Table 5 and a constant 2% porosity. The data fall outside of the generalized field for MORB to more depleted compositions. Batch melting to 30% defines a trajectory terminating in the MORB field, but open-system melting with a 2 wt% melt fraction present in the source (which may be thought of as a constant mantle porosity) produces a trend shown by the solid line in Figure 10. Even if a two-stage batch melting process is envisioned (dashed curve in Fig. 10), with an initial stage of 15% melting followed by melt removal and then an additional 35% batch melting of the residue for a total of 50% melting, the most extreme inclusion compositions cannot be produced. Thermal conditions in a normal mid-ocean spreading environment do not permit melting much beyond a total of 25%-30% due to the increase in solidus temperature when clinopyroxene disappears from the assemblage.

Melting of a peridotite mantle will produce magmas varying in composition with percent melting as well as with pressure (Kushiro, 1969, 1972; Jaques and Green, 1980; Takahashi and Kushiro, 1983; Falloon and Green, 1987). With increasing pressure, melts become less silica-rich, but conform to pseudocotectic curves in the CMAS projections that are subparallel to each other. In a general sense, the

Table 5. Parameters used in melting model.

	Срх	Ol	Opx	Sp	Co
DTi	0.36	0.008	0.08	0.07	872
D7.	0.15	0.003	0.02	0.04	7.1
D ₁	0.05	0.00001	0.008	0.0008	1.0
DCe	0.1	0.00005	0.01	0.0008	1.2
DNA	0.17	0.0001	0.02	0.0008	1.6
Dsm	0.28	0.0007	0.03	0.0009	2.0
DEn	0.34	0.001	0.04	0.009	2.1
DDy	0.43	0.004	0.06	0.0015	2.2
Dur	0.42	0.007	0.1	0.0045	2.3
Dyn	0.42	0.009	0.12	0.0045	2.3
X	0.16	0.56	0.24	0.04	
D	0.67	-0.10	0.38	0.05	

Notes: Cpx = clinopyroxeene; Ol = olivine; Opx = orthopyroxene Sp = spinel; C₀ = starting source composition. X = starting mode in the source modified from Johnson et al. (1990). p = proportions of minerals entering melt derived using experimental data from Takahashi and Kushiro (1983), Falloon and Green (1987, 1988), Falloon et al. (1988), and Kinzler and Grove (1992). Sources of partitioning data: Clinopyroxene: Johnson and Kinzler (1989), Hart and Dunn (1993), K. Johnson (unpubl. data). Olivine, opx: REE, Stosch [1982], combined with cpx partition coefficients, McKay (1986). Ti, Zr, Sr, Fujimaki et al. (1984), Weill and McKay (1975), McKay and Weill (1977). Spinel: REE, Stosch (1982), combined with cpx partition coefficients. Ti, Zr, K. Johnson, Ti calculated from electron probe data; Zr estimated from secondary ion counts relative to cpx. Garnet: REE, Shimizu and Kushiro (1975), Nicholls and Harris (1980), Fujimaki et al. (1984). Ti, Zr, Sr, Fujimaki et al. (1984).

data in Figure 7 define a trend that intersects the isobars from 10 to 5 kilobars. We interpret this trend as representing polybaric melting and entrapment of individual instantaneous melt fractions during magma ascent beneath the ocean ridge. A similar trend was identified in a suite of melt inclusions from the North Atlantic Ocean by Sobolev and Shimizu (1993).

SUMMARY AND CONCLUSIONS

The data and melting model presented suggest that the melt inclusions in plagioclase from diabases in hole 504B represent trapped primitive or primary magmas. This is supported by Mg numbers that could have been in equilibrium with mantle olivines and by trace element models indicating late-stage derivation by open-system melting of a slightly REE-depleted peridotite source with about 2% melt retained in the source. These magmas were formed during pressure-release melting of the mantle over a range of depths between 30 and 15 km and represent small melt fractions produced by a near-fractional, open-system melting process.

The data suggest that magma evolution must be a complex process that starts as soon as melts segregate from their source. We speculate that mixing between melt increments from different sources can occur in channels, conduits, or veins at any point in the ascending mantle after segregation takes place, but that equilibration between these melt increments and the surrounding residual mantle must be minimal. Thus, extremely depleted compositions can be trapped as melt inclusions in growing mineral phases before homogenization of the diverse melt fractions prior to eruption. This study demonstrates that silicate melt inclusions can be useful in gaining insights into melting processes that complement those provided by abyssal peridotites.

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Date of initial receipt: 28 May 1993 Date of acceptance: 14 April 1994 Ms 137/140SR-004